AMIGA at the Pierre Auger Observatory: The interface and control electronics of the first prototype muon counters

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13 Abstract

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AMIGA is an enhancement of the Pierre Auger Observatory. The main goals of AMIGA are to extend the full efficiency range to lower energies of the Observatory and to measure the muon content of extensive air showers. Currently, it consists of 61 detector pairs, each one composed of a surface water-Cherenkov detector and a buried muon counter. Prototypes of the muon counter - buried at a depth of 2.25 m - were installed at each vertex of a hexagon and at its center with 750 m spacing. Each prototype has a detection area of 10 m^2 segmented in 64 scintillation strips and coupled to a multi-anode PMT through optical fibers. The electronic systems of these prototypes are accessible via a service tube. An electronics interface and control board were designed to extract the data from the counter and to provide a remote control of the system. This article presents the design of the interface and control board and the results and performance during the first AMIGA acquisition period in 2012.

- ¹⁴ Keywords: Underground Detector, Segmented Scintillators, Data handling,
- ¹⁵ Detector control systems, Data acquisition concepts

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16 1. Introduction

The Pierre Auger Observatory, optimized for the highest energies of the 17 cosmic ray spectrum, has already studied two cosmic ray spectral features: the 18 ankle and the GZK-cutoff [1][2][3]. However, the cosmic ray energy spectrum has 19 two other observed features at lower energies where the spectral index changes: 20 the knee ($\approx 8 \times 10^{15} \,\mathrm{eV}$) and a second knee ($\approx 8 \times 10^{16} \,\mathrm{eV}$)[4]. The transition 21 from galactic to extragalactic sources is supposed to occur according to models 22 is either in the region near the second knee or along the ankle [5][6][7]. A 23 way to identify this transition would be to measure a change in the cosmic 24 ray composition from dominant heavy primaries to either a mixed or a light-25 dominated composition. Although galactic magnetic fields deflect the particle 26 trajectories, making it impossible to identify the sources in the range of the knee 27 and second knee, composition studies should help to discriminate whether the 28 sources are galactic or extragalactic, and where the transition occurs. 29

The Pierre Auger Observatory has two kinds of detectors, water-Cherenkov 30 detectors and fluorescence telescopes. Enhancements to the Observatory lower 31 the full efficiency range down to the second knee. The fluorescence telescope 32 enhancement is called the High Elevation Auger Telescopes (HEAT)[8]. Addi-33 tionally the surface detector enhancements consist of an infilled area of standard 34 water-Cherenkov detectors deployed in a triangular grid of 750 m spacing, each 35 with an associated muon counter. This latter enhancement is called Auger Muon 36 Infill for the Ground Array (AMIGA). 37

Prototypes of the muon counter were developed and installed in an area designated as the Pre-Unitary Cell (Figure 1). These prototypes consist of a 10 m² scintillation detector segmented in 64 strips. Each strip is 400 cm long, 41 4.1 cm wide and 1 cm thick and made out of extruded polystyrene doped with fluorine and co-extruded with a TiO₂ reflective coating. Strips are placed in two groups of 32 at each side of a central dome where the photomultiplier and electronics are located. Saint-Gobain 1.2 mm diameter optical wavelengthshifting fibers are attached with optical cement to the strips in a groove along the

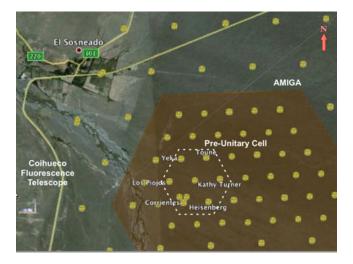


Figure 1: Map of the area for deployment of AMIGA muon counters (brown color). The dotted line encloses the hexagonal array where pairs of water-cherenkov and muon counters have been installed for first prototype tests. The Coihueco fluorescence telescopes and the HEAT extension are located about 5 km to the west.



Figure 2: Two muon counters in the laboratory without their top PVC cover. In the middle of the counters are the ends of the 64 optical fibers and the connector to couple the fiber ends to the PMT.

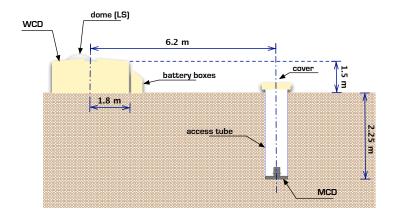


Figure 3: Cross-section of the water-Cherenkov detector (WCD) and the muon counter (MCD) on site.

strips. The fibers end at an optical connector, which is attached and aligned to
a 64 channel multi-anode photomultiplier tube (PMT) H8804 with 2 mm×2 mm
pixels (Figure 2).

To avoid contamination from the electromagnetic component of the shower, the modules are buried underground at a depth of 2.25 m (equivalent to 540 g/cm² of mass overburden of the site soil).

The shielding of the soil imposes a threshold of around 920 MeV[9] for vertical muons and assures negligible electronic contamination because it contains more than $20X_0$ (where X_0 is the radiation length).

As a consequence of the installation of the counters underground, the electronics is split into two components: the underground electronics, which is integrated into the buried muon counter modules and accessible through a service tube, and the surface electronics placed next to the electronics of the water Cherenkov Detector (WCD) (see Figure 3).

The underground electronics includes a 64-pixel PMT, an analog front-end, an FPGA, the interface and control board and a power distribution board (Figure 4). The front-end includes a high voltage power supply, a PMT socket and two low drop-out linear regulators to power the amplifiers for the 64 analog channels of the front end. The power distribution board has separate switching

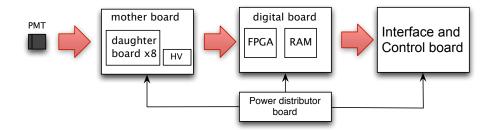


Figure 4: The set of boards implemented in the muon counter that are accessible via the service tube. The prototype version of the electronics is split into a motherboard (plus daughter cards), a digital board, an interface and control board and a power distribution board. The final design will integrate the electronics into only two boards: a front-end and a digital acquisition board.

power supplies to feed the complete set of boards. The electrical ground of both
the buried and the surface electronics is decoupled by floating power supplies
to avoid ground loops and noise.

The output currents of the PMT are converted into voltage pulses and compared with a threshold level. The threshold level can be set individually for each channel by a 12 bit digital-to-analog converter (DAC). Below threshold level, the digital output signal is set to one, while above threshold, the signal is set to zero (inverted logic). The DAC can be programmed by a serial peripheral interface (SPI).

At each trigger, 256 words (64 bits length) from a circular buffer are captured and stored together with additional 512 words from a linear buffer into an external RAM memory. Each time data are requested, the external RAM memory content is transferred to the interface microcontroller (μ C) by a parallel bus.

The mean power consumption of the underground electronics including the
interface and control board is 5.28 W with a peak around 5.52 W.

The surface electronics of the muon counter is located inside the dome of the water-Cherenkov detector. It has two main components: a wireless communication system (TS7260 Single Board Computer from Technologic Systems) and synchronization hardware connected to the electronics of the water-Cherenkov
detector. Additionally, a TSCAN1 board from Technologic System was added
to implement a CAN bus in the counter to handle the data stream between
underground and surface electronics.

88 2. Trigger and acquisition modes

The muon counter can use either an external trigger signal provided by 89 the water-Cherenkov detector or a stand-alone trigger (a coincidence trigger 90 generated with the coincidence of one or more channels). In external trigger 91 mode, the underground electronics receives a trigger pulse (T1[10]) and a local timestamp from the electronics of the water-Cherenkov detector (Local Station 93 or LS). Meanwhile the surface electronics receives a GPS timestamp with the 94 local timestamp from the Local Station. The GPS timestamps and the local 95 timestamps are transmitted from the LS to the Single Board Computer (SBC). 96 Thus, each event recorded by the muon counter is synchronized with the water-97 Cherenkov detector event at T1 level. The latency between the local timestamp 98 and the trigger of the underground electronics is a fixed number given by the 99 delay of the T1 pulse through the cable between the LS and the underground 100 electronics (about a few 10 ns). For the muon counter there is no latency between 101 the GPS timestamp and the local timestamp since they are transmitted at the 102 same time from the LS to the SBC. 103

The memory size of the underground electronics is designed to store 2048 events. At an average T1 trigger rate of 100 Hz this corresponds to a storage time of about 20 seconds.

The counter can be programmed with three acquisition modes implemented in the interface and control board: stand-alone trigger, an external trigger and a calibration trigger. In stand-alone trigger mode, the trigger condition is the presence of a signal in specific channels. The aim of this acquisition mode is to record the trigger rate of a channel or group of channels given a known threshold level and PMT voltage. The external trigger mode is used to store the signal traces when a T1 signal is received by the counter. In the first counter prototypes, the traces have $9.6 \,\mu s$ length with a trigger point at $3 \,\mu s$. Each event is a collection of 768 words of 64 bits corresponding to the 64 detector channels stored at 80 MHz.

All the T1 events are stored locally in the underground electronics until a T3[10] is received from the Central Data Acquisition (CDAS) of the Pierre Auger Observatory.

In calibration mode, the counter operates in stand-alone trigger mode but 120 the threshold level is swept over a range to record the threshold dependent 121 trigger rate per channel. In this case the recorded data are rates and traces. 122 The main goal of this trigger mode is to record the threshold level at a given 123 rate. The data readout in calibration mode is done by programming a fixed 124 acquisition time. The counter records data for the selected time period and 125 then transfers the data to the SBC. The main difference between the external 126 trigger and the other two modes is that with the external trigger the counter 127 has to work in real time. Therefore, in external trigger mode the event data 128 transfer from the underground to the surface electronics has to be faster than 129 the T3 rate in order to avoid data losses. 130

¹³¹ 3. Interface and control board design

The acquisition modes described above define the main interface and control board design requirements. These requirements are described below.

• Physical line: A 20 m cable length (Figure 3) is needed because the mini-134 mum distance between the water-Cherenkov detector and the underground 135 electronics is about 12 m and we added an extra 8 m to have enough flex-136 ibility during cabling or land preparation. The transceiver in the under-137 ground electronics must work with 3.3 V and the transceiver in the surface 138 electronics must work with 5 V due to compatibility reasons between the 139 underground and surface electronics. There are 6 physical lines: two lines 140 for data, two lines for triggering and two lines for power. 141

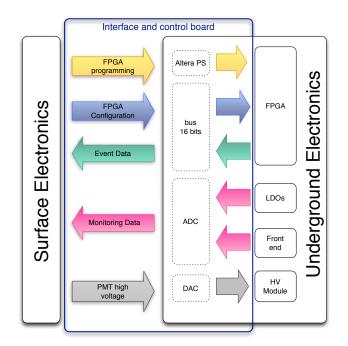


Figure 5: Functional blocks and data flow of the interface and control board. The design of the interface and control board met the acquisition requirements.

• Event transfer rate: The raw event size is 6 kB. The T3 event rate was 142 measured to be 150 events per day (peak value) during a test period of 143 six months. However, the maximum payload occurs while working in 144 calibration mode. This is not real time. The payload in calibration mode 145 depends on threshold levels and acquisition period (i.e. if the threshold 146 level is set below the noise level and the acquisition period is too large, 147 the event rate will be too high). The external memories of the FPGA can 148 only store 2048 events. In this case 1 MB has to be transferred, but there 149 is no time constraint since the counter is working off line. 150

• FPGA programming: This task has to be done each time the system is powered up or in case a new upgrade of the FPGA code is ready to be downloaded. The programming has to be done through a serial connection using the passive serial protocol from Altera.

• FPGA configuration and data handling: FPGA configuration and data 155 request handshake is defined by a 16 bit parallel bus and an 8 bit address 156 bus (with three control signals). A parallel bus was selected to map the FPGA registers in the μ C memory. In this way any access from the μ C to 158 the FPGA is done using standard read and write transactions. The final 159 design will have a 16 bit address bus. 160

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• Slow control: The board has to include a DAC to set the PMT with a 161 gain of 7.7×10^6 (which corresponds to 950 V) and it has to be monitored. 162

• Monitoring: The regulator outputs from the motherboard and the digital 163 board have to be monitored to detect any failure during acquisition. Also 164 the PMT and interface temperatures have to be acquired for the same 165 reason. Additionally, the lines coming from the motherboard have to be 166 isolated from the interface and control board ground. 167

The functionality of the interface and control board is not integrated in the 168 digital board at the prototype stage to allow for parallel development. In the 169 final version, however, all the analog boards will be integrated in a single front-170 end. Also, all the digital boards will be integrated into a single acquisition 171 board. 172

The interface and control board design provides a solution to the require-173 ments: the communication between the underground and the surface electronics, 174 the monitoring and the control of the underground electronics and as an inter-175 face for the automatic processes implemented in the SBC at the surface (Figure 176 5). 177

One of the main design criteria was to include the minimum amount of 178 hardware and to use standard communication protocols. The core of this board 179 is a 32 bit ARM (TMS470) μ C from Texas Instruments running at 20 MHz. 180 An ARM architecture microprocessor was selected because of its low current 181 consumption (110 mA@24MHZ). The schematic in Figure 6 shows the modules 182 programmed in the μC . The communication between the underground and the 183

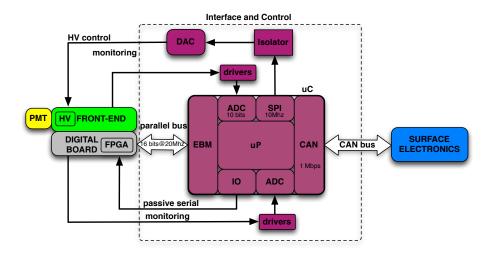


Figure 6: The microcontroller (μC) is in charge of multiplexing data to its corresponding modules. A CAN link between the underground and surface electronics is accomplished by an built-in CAN module.

¹⁸⁴ surface electronics is accomplished by a built-in CAN module because of its low
¹⁸⁵ power consumption and its standard protocol. The data received through the
¹⁸⁶ CAN bus are multiplexed to the corresponding modules.

187 3.1. Physical line: CAN bus

Two constraints related to the maximum length of the physical line were 188 analyzed for the design of the CAN bus: the round-trip delay and the amplitude 189 bit drop, which is the amplitude drop of the analog signals corresponding to 190 one bit. The round-trip delay (RTD) is a critical parameter of the CAN bus 191 concept because the CAN protocol uses a bit-wise arbitration to select which 192 node should continue signalling. Thus the bus length is limited by RTD to avoid 193 bit corruption due to delayed bits being sensed by other nodes. The round-trip 194 delay was calculated for the selected design (Figure 7). As conservative estimate 195 we assume a delay of 215 ns.

¹⁹⁷ Using the equation

$$t_{RTD} = \frac{1}{baudrate} \tag{1}$$

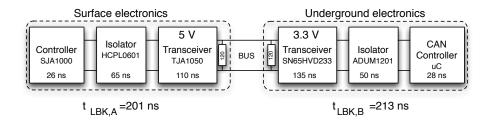


Figure 7: Signal path from surface to underground electronics used in the design.

and after substituting with $t_{RTD} = bus_{prop} \times L_{max}[11]$ it is found that the maximum bus length (L_{max}) at the maximum baudrate of 1 Mbps would be 57 m (using UTP cable CAT5e with $bus_{prop}=5$ ns/m):

$$L_{max} = \frac{baudrate^{-1} - 2 \times t_{loop}}{2 \times bus_{prop}} = 57 \,\mathrm{m} \tag{2}$$

Nevertheless, the maximum achievable bus line length in a CAN bus network is also determined by the amplitude due to the series resistance of the bus cable and the input resistance of the bus nodes. This relationship is expressed in the following equations:

$$L_{max} = \frac{1}{2 \times \rho} \times \left(\frac{V_{diff.out.min}}{V_{th.max} + \Gamma_2} - 1\right) \times \frac{R_{T.min} \times R_{diff.min}}{R_{diff.min} + R_{T.min}}$$
(3)

where

$$\Gamma_2 = k_{sm} \times (V_{diff.out.min} - V_{th.max}) \tag{4}$$

and $\rho =$ is the specific resistance per length unit.

Here k_{sm} is the safety margin expressed as the fraction of the difference between the output level at the transmitting node and the receiver input threshold for detection of a dominant bit [11]. Thus, the maximum cable length is estimated to be 25 m when taking a safety margin of 75% and the worst transceiver type (see Figure 8).

Eye patterns provide a good representation of how the data have been affected by a transmission line. Positive and negative pulses are superimposed on each other. Overlaying many bits produces an "eye" diagram, so called because the resulting image looks like the opening of an eye.

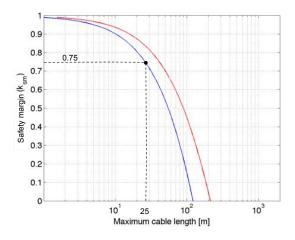


Figure 8: Bus length dependence on the given safety margin. The red and blue lines represent the maximum lengths using underground and surface transceiver parameters, respectively.

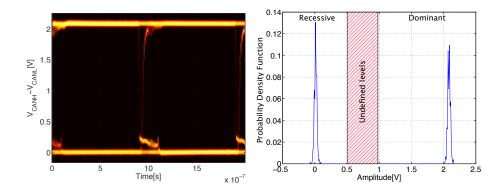


Figure 9: Eye pattern and vertical histogram at sample point (t=1.8 μ s, Number of traces=1×10⁶). Recessive and dominant levels appear well defined at 0 V and 2.1 V. The forbidden area is colored red. In this area, bus levels are undefined for the transceivers.

In the setup, random patterns were generated at the surface electronics and 211 sent by CAN bus to the underground electronics. The patterns were received 212 by the μ C and then echoed to the surface. Since CAN uses a differential bus, 213 the signals were measured in the CANH (high level port of CAN) and CANL 214 (low level port of CAN) ports with an oscilloscope. The measurements were 215 done using the surface transceiver because the transceiver selected for the un-216 derground (SN65HVD233) is powered with 3.3 V. Thus, the most demanding 217 configuration is when the TJA1050 transceiver (5 V) receives signals from the 218 underground transceiver (3.3 V). The results of the tests are shown in Figure 219 9 where one can see the eye pattern and the vertical histogram at the sample 220 point (t=1.8 μ s). The distribution of both levels (dominant and recessive) are 221 well defined with peaks at 0 and $2.1 \,\mathrm{V}$. None of the 10^6 pulses recorded fell in 222 the forbidden area where the bus levels are undefined for the transceivers thus 223 ensuring a proper transmission. 224

225 3.2. Passive serial bus

Passive Serial (PS) is a programming method that can be performed on the 226 Cyclone III device family with an external intelligent host, such as a micropro-227 cessor [12]. In the PS scheme, a μ C controls the configuration of the FPGA. In 228 this mode the configuration data are clocked into the Cyclone III device using 229 the DATA0 pin at each rising edge of the DCLK (clock). A simple routine 230 was programmed into the μC to receive the configuration file from the surface 231 electronics via the CAN bus and then transfer it by PS to the Cyclone using 232 five I/O general purpose ports: nCONFIG, nSTATUS, CONFDONE, DLCK 233 and DATA0. During configuration, the Cyclone III (FPGA) decompresses the 234 bitstream in real time and programs SRAM cells. This decompression feature 235 is supported in PS mode. As mentioned in the Cyclone III handbook [12], 236 compression reduces the configuration bitstream size by 35-55%; thus the data 237 transfer during configuration in the CAN bus is packed with a certain compres-238 sion factor reducing its size and the transfer time. 239

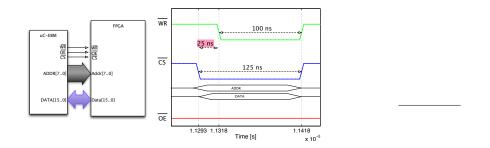


Figure 10: Left:Interconnection between μ C and FPGA using the EBM. Right: EBM signal timing for write cycle: Data were taken with a logic analyzer and then plotted.

240 3.3. Parallel bus

In the underground electronics, the core of the acquisition system is implemented in a Cyclone III (FPGA) and a memory bank for event data storage.

The interconnection between the μC and the FPGA to configure its registers 243 is implemented by a data bus of 16 bits, an address bus of 8 bits and three control 244 signals (see Figure 10). In the μ C, these ports are supported by the Expansion 245 Bus Module (EBM) designed to connect external memories. In this way, the 246 FPGA is directly mapped to the μ C and the FPGA registers can be accessed 247 by the μC as an external memory. Each μC transaction requires a minimum 248 of one clock cycle. Nevertheless, five wait states were added because the EBM 249 is clocked by a 40 MHz internal Phase-Locked Loop (PLL) and the μ C internal 250 PLL is not synchronized with the FPGA clock. This provides the stability 251 required for the desired performance. Using this configuration each access cycle 252 to the FPGA takes 125 ns (Figure 10). 253

The tests were done by performing 10000 write/read cycles into the FPGA registers to detect any perturbation in the synchronization between the FPGA and the μ C (not synchronized with the same clock) during the access to the bus using five wait states. In order to achieve that, random data were generated by the SBC and sent to the underground electronics. The μ C wrote each pattern, read it and sent it back to the SBC where patterns were compared. No data loss was detected using five wait states.

²⁶¹ 3.4. High voltage control

The PMT high voltage (HV) is provided by a module from Hamamatsu 262 (C4900-1). As the HV will never change during normal operation, the control 263 of this parameter is just for compensating the gain due to PMT aging. However 264 it is known that if voltage is applied abruptly to a tube connected in negative 265 polarity, the amplitude of the initial dark-current transient may be high enough 266 to damage the sensitive measuring apparatus. Therefore the μC can apply the 267 voltage gradually to reduce the transient. The controlling voltage input of the 268 C4900-1 varies from 0 V to 5.3 V [13]. The control of the HV module input is 269 performed by the μ C using an external 10 bit DAC that is connected to the 270 μ C via SPI (Serial Peripheral Interface). In this way the μ C can set any value 271 between -3 V and -1000 V with a minimum step of 1.2 V. 272

The SPI application is configured at a rate of 10 MHz with a word length of 16 bits. The DAC (TLV5617A from Texas Instruments) is used with an external reference of 2.5 V. The resistor string output voltage of the TLV5617A is buffered by a ×2 gain rail-to-rail output buffer. The buffer features a Class-AB output stage in order to improve stability and to reduce settling time, and provides an output voltage at full scale given by

$$Vout = \frac{2 \times REF \times CODE}{2^n} [V].$$
(5)

2	7	c

Here REF is the reference voltage of 2.5 V, CODE is the digital input value within the range of 0 to 2^{n} -1 and n = 10 (bits).

As the DAC is powered with a 5V regulator, its internal amplifier is used to get a full range (0-5V) and to improve its output linearity. The 3.3V output from the motherboard is used to power a 5V regulator isolating the digital section of the interface board. Furthermore an isolator (ADuM1400ARW) was added between the μ C SPI and the DAC ports. This configuration allows the μ C to transfer the control voltage codes from the surface electronics directly to the SPI module.

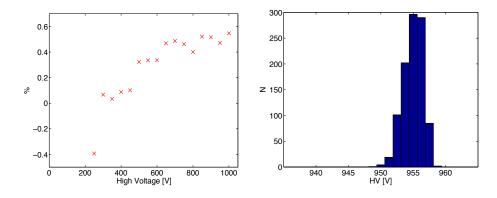


Figure 11: Left: HV was programmed in 50 V steps and measured; the y axis was calculated with $\left(\frac{V_{measure} - V_{fit}}{V_{measure}}\right) \times 100$. Right: PMT HV was set at 954 V; 1000 measurements were taken with the monitoring circuit. The standard deviation observed was ± 1.50 V

A test to check the DAC offset and gain errors was done by programming high voltage values and plotting them versus the measured high voltage values. To determine deviations from a linear behaviour the points were fitted with a linear function $a_1 + a_2 \times HV$ resulting in an offset error a_1 =-2.25 and an gain a_2 =1. The plot of the residuals in Figure 11 confirms a deviation from linearity below 0.6 %. over the full range.

295 3.5. Monitoring

The interface and control board provide twelve 10 bit ADC channels to mea-296 sure parameters of the underground electronics. The channels are used to mon-297 itor the PMT high voltage, the supply voltage of 12 V, 3.3 V, -3.3 V from the 298 front-end regulators, and the 3.3 V, 1.2 V and 2.5 V of the FPGA power supply. 299 In addition, the μ C and PMT temperatures are monitored by sensors from Ana-300 log Devices (AD22103). The sensors provide a voltage level that is digitized by 301 the ADC. The input amplifiers for the ADC channels are included in the inter-302 face and control board. Two schemes are implemented taking into account that 303 the ground reference of the ADC is not the same as the motherboard ground. 304 As an example, one of the circuits implemented to monitor a voltage level 305

³⁰⁶ from the motherboard is shown in Figure 12. The PMT high voltage module

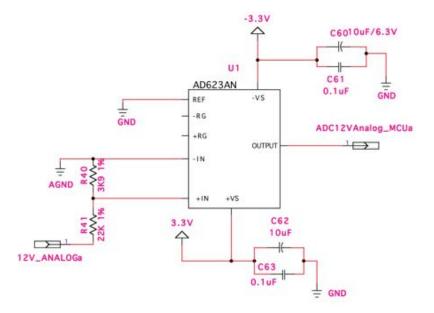


Figure 12: Monitoring circuit implemented for 12 V signal. The signal comes from a regulator to power the high voltage module in the motherboard.

uses a 12 V regulator and it is monitored by the interface board using a resistive 307 divider together with a differential amplifier at unity gain. The HV module 308 reference terminal defines the zero output voltage. This is useful if the load does 309 not share a common ground with the rest of the system as in this application. 310 Because the AD623 output voltage is set with respect to the potential on the 311 reference terminal, the grounding problem is solved by connecting the REF pin 312 to the local ground (GND in Figure 12). Because the motherboard includes a 313 HV monitoring circuit with a scale factor of 1:1000 the same configuration shown 314 before is used to measure the PMT high voltage level. Similar configurations 315 are implemented with $\pm 3.3 \,\mathrm{V}$ regulators. 316

Additionally, the digital board regulators, which do share the ground level with the ADC, are interfaced as shown in Figure 13, where the signals are only buffered and adapted to the ADC input ranges.

A gain around 7.7×10^6 is expected with 954 V and this gain value was selected in the laboratory to get a good SPE resolution. The standard deviation

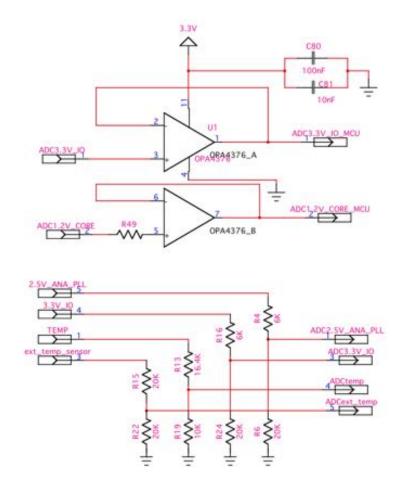


Figure 13: The signals 2.5V_ ANA_ PLL and 3.3V_ IO come from regulators used to power the FPGA. TEMP and ext_ temp_ sensor come from the interface board.

 $_{322}$ obtained is ± 1.50 V which represents a low dispersion.

Also, monitoring measurements were taken in the laboratory with the interface board to check the components stability as a function of temperature in the range of interest. The PMT high voltage was set to 954 V (anode to cathode). The PMT temperature was increased using a heat resistor, since some preliminary tests of the mechanical design of the module showed a daily thermal and a seasonal excursion inside the underground electronics dome[14]. One of the temperature sensors from the interface and control board was attached to the

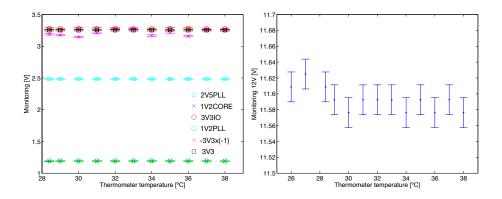


Figure 14: Left: Monitoring of regulator output levels implemented in the underground electronics. Output levels seem to be stable in the whole temperature range. The negative power supply (-3.3 V) shows fluctuations that could be related to bad filtering. Right: Monitoring of the 12 V regulator used to power the high voltage module. Error bars show the accuracy of the measurements. The deviation from the nominal value of 12 V was detected at room temperature. It shows that the regulator was working outside the component specification (output voltage accuracy of 1 %).

PMT socket. Meanwhile, a second sensor was added to the interface and control board. Both these sensors provided the voltages to the ADC. A reference thermometer (with a type K thermocouple) was located inside the electronics enclosure next to the sensor located in the interface and control board. Since both sensors were located next each other, a difference was not expected in the measurements of the distance to the source of heat. The monitoring system acquired the temperature measurements and all the monitoring values.

The results of temperature stability in the acquisition chain of the monitoring 337 signals are plotted on the left side of Figure 15. As can be seen from the figure, 338 there is only a negligible dependence on temperature. The comparison of the 339 reference temperature and the integrated temperature sensor is shown on the 340 right side of Figure 15. The measurements are in agreement, but the integrated 341 sensor shows a bias of about 2°C, which is within the specified absolute accuracy 342 and which can be corrected by a temperature calibration. The error bars on 343 the right side of Figure 15 are the errors provided by the manufacturer of the 344 reference thermometer and the errors due to the circuits implemented in the 345

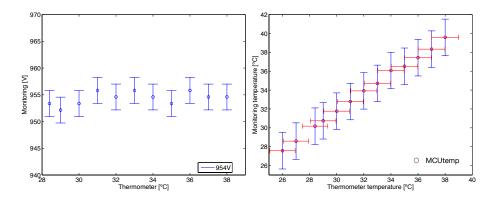


Figure 15: Left: Monitoring of the high voltage set to the PMT. The DAC was programmed to set 954 V in the HV module. The selected HV value corresponds to the selected gain for the PMT. Right: Temperature comparison between thermometer measurements and interface sensor. A systematic error is found and could be improved with a temperature calibration.

³⁴⁶ interface and control board.

Monitoring data are transmitted to the central data acquisition by request. The idea of the monitoring is to check the voltages around a nominal value. If voltages are outside the manufacturer specification range, corrective actions have to be taken by a monitoring control. Implementation of a monitoring central is foreseen with alarms management.

³⁵² 4. Acquisition and compression algorithm

The Pre-unitary Cell was the first array of counter prototypes deployed in AMIGA and data were recorded during one year (2012). The footprint of a shower with the core falling within the prototype hexagon is shown in Figure 16 along side the digital trace patterns recorded by the muon counters participating in this event.

Basically, data in the buried counters are active samples within a region around the trigger bin 256 (length of the circular buffer). The spread around the trigger bin of the active samples provides a rough estimate of the time width of the shower front. An active sample represents one sample of the input signal having an amplitude above the threshold level. The result of a three-month

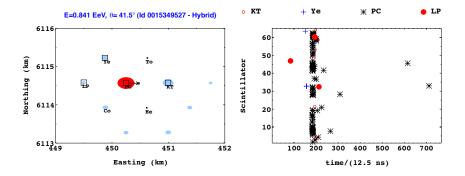


Figure 16: Example of a T3 shower event. LP, Ye, To, KT, He, Co and PC are detector pairs. Left: Pair positions, the red point represents the shower impact, sky-blue points mark Auger detectors triggered in the event. Square points are the muon counters which recorded data. Right: Scintillator strip positions are represented in y axis. Samples in time are represented in x axis. PC was the Auger detector with the highest signal. LP, Ye and KT were also triggered.

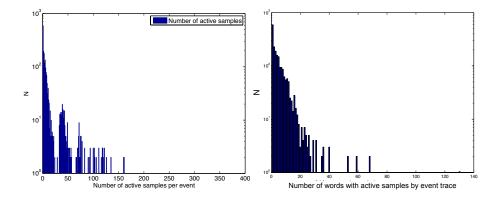


Figure 17: Left: Distribution of the number of active samples for T3 events in 3 months of data from the Pre-Unitary Cell. The distribution is not continuous, and there are some groups. The analysis was made only to explain data compression. Groups could be related to trace shapes. Right: Number of words with active samples. Events collected with the Pre-Unitary Cell with a threshold level of 200 mV.

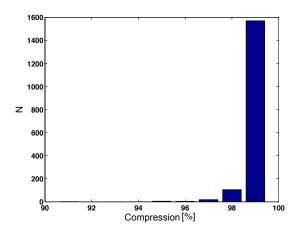


Figure 18: Compression applied to data shown in the right side of Figure 17. The compression of data traces is bigger than 90%.

period of data analysis from the Pre-Unitary Cell is that the maximum number 363 of active samples in a T3 event is 381 and most of the active samples cease at 364 approximately 160 as shown in Figure 17. The maximum number of samples 365 (381) represents only 0.77% of the maximum possible value of 49152 (768 \times 366 64), where all the samples of an event are in an active state. Later, a simple 367 compression algorithm was implemented in the μC of the interface and control 368 board to improve the data transfer up to the surface electronics. Only time bins 369 of event traces from channels with active samples are transferred along with its 370 corresponding bin position. 371

The results of the compression algorithm can be seen in Figure 17. The minimum compression achieved is 90% while the average compression is 98.8%.

374 5. Results and discussion

Studies and measurements of the interface and control board functions (data transfer, FPGA programming, monitoring and slow control) and CAN bus characteristics were performed and analyzed. It was found that the CAN protocol is a suitable solution for a single AMIGA module or more than one interconnected module through a CAN bus (backbone topology) having a cable length of 25 m

(considering a safety margin of 75%). This system has been successfully imple-380 mented in the Pre-Unitary Cell construction phase of the AMIGA project and 381 particle shower events were successfully acquired and transmitted through the 382 interface and control board. Data recorded with the Pre-Unitary Cell were used 383 to implement a compression algorithm in the μC of the interface and control 384 board to reduced the amount of data transfers. The minimum and maximum 385 compression factor was found to be 90% and 99% respectively. The acquisition 386 mode flexibility of the interface and control board allowed performing several 387 tests with the counters and it was used for engineering re-design purposes. 388

389 6. References

[1] Alexander Schulz for the Pierre Auger Collaboration. The measurement of
 the energy spectrum of cosmic rays above 3 × 10¹⁷ eV with the Pierre Auger
 Observatory, in proceedings of 33rd International Cosmic Ray Conference,
 2013, (Río de Janeiro-Brasil)

- ³⁹⁴ [2] Diego Ravignani for the Pierre Auger Collaboration. Measurement of the ³⁹⁵ energy spectrum of cosmic rays above 3×10^{17} eV using the AMIGA 750 ³⁹⁶ m surface detector array of the Pierre Auger Observatory, in proceedings of ³⁹⁷ 33^{rd} International Cosmic Ray Conference, 2013, (Río de Janeiro-Brasil)
- [3] The Pierre Auger Collaboration, Measurement of the energy spectrum of cos mic rays above 10¹⁸ eV using the Pierre Auger Observatory, Physics Letters
 B 685, (2010) 239-246.
- [4] W. D. Apel et al. (KASCADE-Grande Collaboration), Kneelike Structure
 in the Spectrum of the Heavy Component of Cosmic Rays Observed with
 KASCADE-Grande, Phys. Rev. Lett.107, 171104 (2011).
- [5] D. Allard et al, On the transition from galactic to extragalactic cosmic-rays:
 Spectral and composition features from two opposite scenarios, Astrop. Phys.
 27 (2007) 61-75.

- [6] V. Berezinsky et al, Extragalactic UHE proton spectrum and prediction for
 iron-nuclei flux at 10⁸-10⁹ GeV, Astrop. Phys. **21** (2004) 617-625.
- [7] A. Wolfendale, T. Wibig, At what particle energy do extragalactic cosmic
 rays start to predominate?, J. Phys. G 31 (2005) 255.
- [8] C. Meurer, N. Scharf on behalf of the Pierre Auger Collaboration, *HEAT a low energy enhancement of the Pierre Auger Observatory, Astrophys. Space Sci. Trans.* 7 (2011) 183-186.
- [9] J. Beringer et al. (Particle Data Group) The Review of Particle Physics,
 Phys. Rev. D86, 010001 (2012 and 2013 partial update for the 2014 edition.).
- [10] J. Abraham et al, Trigger and aperture of the surface detector array of
 the Pierre Auger Observatory, Nuclear Instruments and Methods in Physics
 Research A. 613 (2010) 29-39.
- [11] H. Eisele, PCA82C250/251 CAN Transceiver, Philips Semiconductors
 Note. AN96116 (1996).
- ⁴²¹ [12] Altera Corporation., Configuring Altera FPGAs, Altera
 ⁴²² Datasheet. CF51001-3.0 (2012).
- [13] Hamamatsu Photonics K.K., Electron Tube Division. Compact high voltage
 power supply units C4900 series., Hamamatsu Datasheet. TACC1013E04
 (2008).
- [14] F. Suarez, Pierre Auger Observatory surface and underground detectors operation and techniques., Dottorato di Ricerca in Fisica Fondamentale, Applicata ed Astrofisica. Facoltà di Scienze M.F.N. Dipartimento di Fisica
 Generale. Univarsità degli studi di Torino. (XXI ciclo a.a. 2005-2008).