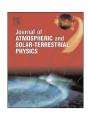
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A new South American network to study the atmospheric electric field and its variations related to geophysical phenomena



J. Tacza ^{a,*}, J.-P. Raulin ^a, E. Macotela ^a, E. Norabuena ^b, G. Fernandez ^c, E. Correia ^{a,d}, M.J. Rycroft ^e, R.G. Harrison ^f

- a Centro de Radio Astronomia e Astrofísica Mackenzie (CRAAM), Escola de Engenharia, Universidade Presbiteriana Mackenzie, São Paulo, SP, Brazil
- ^b Instituto Geofísico del Perú, IGP, Lima, Peru
- ^c Complejo Astronómico El Leoncito, CASLEO, San Juan, Argentina
- ^d Instituto Nacional de Pesquisas Espaciais, Ministério da Ciência e Tecnologia, São José dos Campos, Brazil
- ^e CAESAR Consultancy, 35 Millington Road, Cambridge CB3 9HW, UK
- f Department of Meteorology, University of Reading, Earley Gate, Reading RG6 6BB, UK

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ABSTRACT

In this paper we present the capability of a new network of field mill sensors to monitor the atmospheric electric field at various locations in South America; we also show some early results. The main objective of the new network is to obtain the characteristic Universal Time diurnal curve of the atmospheric electric field in fair weather, known as the Carnegie curve. The Carnegie curve is closely related to the current sources flowing in the Global Atmospheric Electric Circuit so that another goal is the study of this relationship on various time scales (transient/monthly/seasonal/annual). Also, by operating this new network, we may also study departures of the Carnegie curve from its long term average value related to various solar, geophysical and atmospheric phenomena such as the solar cycle, solar flares and energetic charged particles, galactic cosmic rays, seismic activity and specific meteorological events. We then expect to have a better understanding of the influence of these phenomena on the Global Atmospheric Electric Circuit and its time-varying behavior.

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1. Introduction

The confirmation in 1753 (Parsons and Mazeas, 1753; Canton, 1753) that the atmospheric electric field recently identified in thunderstorm conditions persists in fair weather immediately raised questions of how it originates and by which mechanism (s) it is maintained. Here fair weather conditions means that no local electrification processes are occurring in the atmosphere, and the absence of appreciable convective clouds.

Investigations of fair weather electrification phenomena were performed during the late eighteenth and nineteenth centuries, but major advances came with measurements made by the Carnegie Institution of Washington in the early twentieth century. On board a ship, the vertical atmospheric electric field in the atmosphere was measured every hour in the oceanic air. Analysis of the results showed that the daily electric field variation with Universal Time was independent of the ship's position. This diurnal variation is a key part of our contemporary understanding

of the Global Atmospheric Electric Circuit (GAEC), and is generally known as the Carnegie curve (Harrison, 2013). The Carnegie curve was obtained from the averaging of thousands of measurements of the Earth's fair weather electric field in terms of the Universal Time of the measurements over the world's oceans where the planetary boundary layer is relatively free of pollution. In fair weather, the vertical atmospheric electric field is about $E=-130~\rm V/m$, where the minus sign indicates that the electric field vector is pointing downwards. Meteorologists refer to the potential gradient (PG) rather than the electric field (E). The PG and E have the same magnitude, although by convention the PG is positive in fair weather.

The GAEC is formed between the Earth's surface, which is a good conductor of electricity, and the ionosphere, a weakly ionized plasma at and above $\sim\!80$ km altitude. Between these two layers the atmosphere is a reasonably good electrical insulator, i.e. it is a leaky dielectric medium. Electrical "batteries" exist below or inside electrified clouds (e.g., precipitation as rain or thunderclouds) to maintain the atmospheric electric field, implying that an electric current flows up to the ionosphere above these clouds. The electric circuit is closed by downward currents flowing through the fair weather and semi-fair weather regions of the Earth's atmosphere,

^{*} Corresponding author. Fax: +55 11 32142300. E-mail address: josect1986@gmail.com (J. Tacza).

and through the rocks and oceans of the Earth's surface (Rycroft et al., 2008).

The well-known close correlation between the electrical generating system of the GAEC and the Carnegie curve has been mentioned by several authors (Rycroft et al., 2000, 2008; Williams, 2009; Harrison, 2013). However, the exact link is not well understood and continuous monitoring of the atmospheric electric field is required in different geographical areas to help in the interpretation of the behavior of the GAEC (Rycroft et al., 2012).

Solar and space weather phenomena affect the ionosphere causing disturbances in the GAEC (Rycroft et al., 2000, 2012), which can be monitored by analyzing the variations of the atmospheric electric field in fair weather. Transient variations of the intensity of the atmospheric electric field have also been systematically reported in association with time fluctuations of the secondary cosmic ray flux measured at the same location (De Mendonça et al., 2010). Such observations open up new possibilities for a better understanding of the relationship between extreme variations of the electrical properties of the atmosphere, such as during lightning, and Forbush effect decreases.

The GAEC has also been studied to look for short-term precursors of seismic activity. The main reason is that increased radon emanation and/or electrical charge generated by cracking rocks and associated in time and space with strong seismic events is nowadays an undisputed observational fact (see, for example, Ghosh et al. (2009) and Pulinets and Boyarchuk (2004)). Furthermore, the radon decay process dominates the amount of atmospheric ionization in the near-Earth boundary layer and, therefore, changes of radon concentration there will significantly affect the electrical conductivity, which in turn will produce changes on the measured atmospheric electric field (Pierce, 1976). However, these changes are difficult to observe because of the great variety of phenomena and perturbations that can affect its value, such as fog. strong winds, precipitation, aerosols, and pollution (Bennett and Harrison, 2007). Therefore, the search for earthquake precursor effects in the atmospheric PG (Harrison et al., 2010, 2014) needs the behavior of the atmospheric electric field in fair-weather conditions to be known for various meteorological variable conditions.

2. Scientific objectives of the network

In this section we describe the scientific motivations and objectives for the operation of a new atmospheric electric field monitoring network in the South American region.

2.1. Relationship between the Carnegie curve and the generators of the Global Atmospheric Electric Circuit

According to the classical picture of atmospheric electricity, the overall thunderstorm activity at any time charges the ionosphere to a positive potential of several hundred thousand volts with respect to the Earth's surface. This potential difference drives a vertical electric conduction current downward from the ionosphere to the ground in all fair-weather regions of the globe (Rycroft and Harrison, 2012; Rycroft et al., 2000). Horizontal currents flow freely within the highly conducting Earth's surface and in the ionosphere. A current flows upward from the top of thunderstorm clouds toward the ionosphere and also from the ground into the thunderstorm generator, closing the circuit (Rycroft et al., 2000). At first, thunderstorms were thought to be the only "batteries" that maintain the global electric circuit. This was shown in the study of Whipple (1929) which compared the Carnegie curve with variations of the world's thunderstorm areas. However, these two differ quantitatively in two key aspects: a) the relative amplitude variation of the thunderstorm areas over land is more than twice that of the Carnegie curve, and b) the time of maximum global thunderstorm activity corresponds to afternoon local times in Africa, whereas the Carnegie curve maximizes for afternoon local times in South America (Liu et al., 2010).

In order to try to understand these discrepancies, Kartalev et al. (2006) used a theoretical model to estimate the effects of upward currents from a thunderstorm to the ionosphere. They showed that the Carnegie curve preferentially reflects the longitudinal distribution of thunderstorms within 11° magnetic latitude of the magnetic dip equator. South American thunderstorms are close to the magnetic dip equator, whereas most African thunderstorms occur over the Congo at a more southerly dip latitude between 12° and 20°. Thus, South American thunderstorm sources seem to have a greater influence on the Carnegie curve than the more numerous thunderstorm sources over Africa.

In a more recent study, Liu et al. (2010) compared the Carnegie curve with the Universal Time diurnal variations of rainfall from thunderstorms and electrified shower clouds (ESCs) using a 10year database from the Tropical Rainfall Measuring Mission (TRMM). These authors concluded that not only do thunderstorms contribute to the Carnegie curve, but rainfall outside thunderstorm regions is also important. The same conclusions were reached for fair weather electric field measurements made in Vostok, Antarctica, performed by Burns et al. (2005). Blakeslee et al. (2014) using data obtained from the Optical Transient Detector (OTD) and the Lightning Imaging Sensor (LIS) satellites, complemented with other electrical observations from high altitude aircraft missions, found that the total global current is distributed as follows: Land thunderstorms (52%), ocean thunderstorms (31%), non-lightning producing ocean ESCs (15%) and land ESCs (2%). These findings are consistent with predictions by Wilson (1903) that negative charge is carried down by rain (precipitation) and that this current generator plays a part in maintaining the global electrical circuit. Similarly, Wilson (1920) speculated that both thunderstorms and electrified shower clouds account for the descent of negative charge. Despite Africa's dominance in lightning occurrence, its contribution to the diurnal variation of the global circuit is less important than that of the Americas because of the larger electrified rainfall contribution from this region.

It is now known that the Global Atmospheric Electric Circuit, driven by thunderstorms and electrified rain/shower clouds, is also influenced by energetic charged particles from space (Rycroft et al., 2012). However, more observational studies in different geographical regions are needed to investigate this in detail. For this reason, one of the objectives of the new network is to provide reliable electrical field time profiles in fair weather conditions to help obtain a better understanding of the behavior of the GAEC.

2.2. Solar-terrestrial effects

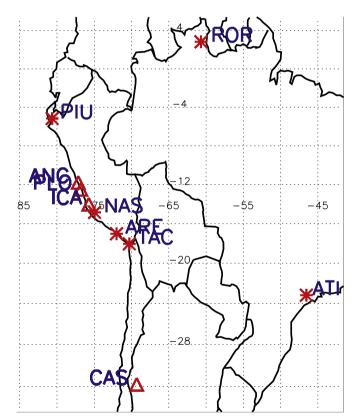
Solar activity varies over a cycle with a period of \sim 11 years. The solar magnetic field and its extension through the interplanetary medium up to the Earth's orbit and beyond is much more variable at solar maximum than at solar minimum. The flux of galactic cosmic rays and the flux of energetic charged particles from the Sun are also subject to this variation. These energetic charged particles as well as those precipitating from the magnetospheric radiation belts during geomagnetic disturbances interact with the Earth's middle and lower atmosphere. They deposit their energy by ionization, by modifying atmospheric chemistry, or by affecting the nucleation of water droplets to form clouds (Rycroft et al., 2000). For these reasons the solar cycle plays an important role in the variations of the GAEC. In addition to long-term changes (i.e. solar cycle variations), the global electric circuit is also affected by short-term solar variations, such as solar flares and solar magnetic

sector boundary crossings in interplanetary space. A summary of how the global electric circuit is affected by long-and short-term solar variations can be found in Rycroft et al. (2012). Although there is some evidence for a solar influence on atmospheric electrical parameters, the solar-terrestrial electrical coupling mechanisms are not well understood (Rycroft et al., 2012).

2.3. Cosmic rays

The fluxes of galactic and solar cosmic rays entering the Earth's atmosphere are essential to the GAEC, since these charged particles are the main source of atmospheric ionization at about 15 to 20 km altitude, and above (Rycroft et al., 2012). The propagation through the atmosphere of secondary cosmic rays can in principle be affected by the presence of electric fields (Muraki et al., 2004; De Mendonca, 2011).

Markson (1981) found positive correlations between variations of the cosmic radiation measured on the ground, and changes of the ionospheric potential derived from aircraft and balloon measurements. Ground measurements show cosmic ray flux excesses



 $\pmb{\text{Fig. 1.}}$ EFM sensors already installed (triangles), and future locations for new sensors (stars).

Table 1Description of existing sensors EFM stations, installed in South America.

in the presence of electric fields associated with thunderclouds (Muraki et al., 2004; De Mendonça et al., 2009; Raulin et al., 2014). Moreover, recently Toropov et al. (2013) found that neutron bursts can be produced in association with negative lightning discharges. These results indicate that cosmic ray showers in the atmosphere may play an important role in the physical processes responsible for cloud formation and lightning discharges.

Data from the new network should lead to a better understanding of the relation between changes in the electrical properties of the atmosphere and measured changes in the simultaneously observed cosmic ray fluxes.

2.4. Atmospheric electric field and seismic activity

The variation of the atmospheric electric field in fair-weather has been analyzed to look for short-term precursors of seismic activity. One of the main reasons for this is the close association between the emanation of radon gas and the occurrence of seismic activity (see the review by Ghosh et al. (2009)). The relationship between radon emanation and the atmospheric electric field has been studied qualitatively by Pierce (1976) who indicated that the release of radon from the ground may increase substantially before an earthquake, causing a decrease in the atmospheric electric field in fair-weather. Recent studies (Harrison et al., 2014) have proposed a mechanism for modeling ionospheric disturbances, by considering Atmospheric Lithosphere-Ionosphere Charge Exchange (ALICE) processes during radon gas emanations (see, for example, Pulinets and Boyarchuk (2004) and Harrison et al. (2010)). These effects are evidenced quantitatively in Hao et al. (1998) and Silva et al. (2012). Finally, Mikhailov et al. (2004) reported variations of the fair-weather atmospheric electric field power spectrum before an earthquake. For these reasons, the new network has been strategically installed in highly seismic areas, one of its objectives being the search for atmospheric electric field anomalies prior to seismic events.

2.5. Meteorological variations

Atmospheric electricity and its relationship with meteorological phenomena have been studied since the eighteenth century. Lord Kelvin at this time had no doubt that the study of atmospheric electrical indicators would bring important clues to increase our knowledge of weather forecasting (Thomson, 1872). Observations confirm that weather changes (fog, wind direction, rain, hail, sleet, snow, etc.) are often associated with pronounced changes of the mean atmospheric electric field values in fair weather (Dolezalek, 1963; Nizamuddin and Ramanadham, 1983; Bennett and Harrison, 2007).

Thus, one of the objectives of the new network is to collect data and to continue the study of fair-weather atmospheric electric field measurements associated with the variations of meteorological parameters. Therefore, in principle, one should be able to

| Location | Latitude (deg) | Longitude (deg) | Height (m) | Description |
|----------|----------------|-----------------|------------|--|
| CAS1 | -31. 799 | -69.297 | 2480 | Installation in 01/2008. The sensor is located away of housing, surrounded scrub and low bushes in a mountainous region and away from industrial areas. |
| CAS2 | -31.800 | -69.293 | 2480 | Installation in 01/2010. Similar characteristics to CAS1. |
| ICA | - 14.089 | -75.736 | 402 | Installation in $12/2011$. The sensor is located in the center of the city of Ica away of housing. Some trees at a distance of 20 m . |
| PLO | -12.504 | -76.798 | 85 | Installation in 11/2011. The sensor is placed on the roof of housing, about 220 m away from the sea and 1700 m from the nearest town. No presence of nearby vegetation. |
| ANC | - 11.777 | - 77.151 | 51 | Installation in $10/2012$. The sensor is placed on the roof of housing, about 1000 m of a highway and 1200 m from the nearest town, and about 2500 m away from the sea. Few trees at a distance of 40 m . |

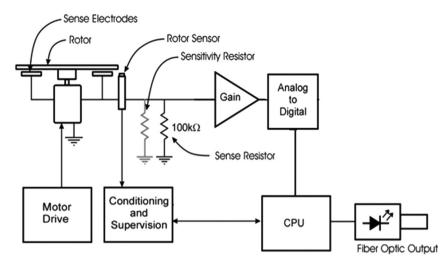


Fig. 2. Block diagram for EFM sensor (from Boltek Corporation EFM100-1000120-050205).



Fig. 3. Final position of the EFM sensor installed at El Leoncito (Argentina).

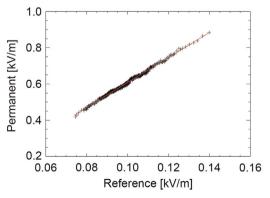
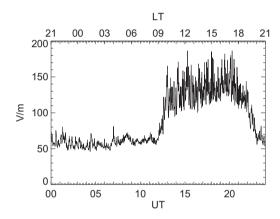


Fig. 4. Linear regression of EFM measurements, obtained at El Leoncito (Argentina), at ground level (Reference) and at the permanent final position (Permanent). Then, the values are corrected as follows: E permanent=mE reference+c, where m=7.36 and c= -0.09 kV/m.



 $\textbf{Fig. 5.} \ \, \text{Diurnal variation of the atmospheric electric field in fair-weather on 31 August, 2011.}$

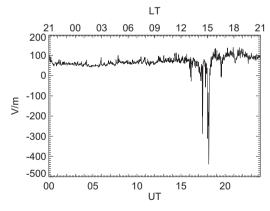


Fig. 6. Diurnal variation of atmospheric electric field in disturbed-weather on 26 August, 2011.

infer whether the electric parameters are only passive symptoms or active partners in these atmospheric processes.

3. Installation of the new network, data processing and first results

Continuous measurements of the atmospheric electric field are being recorded at various locations in South America. The already

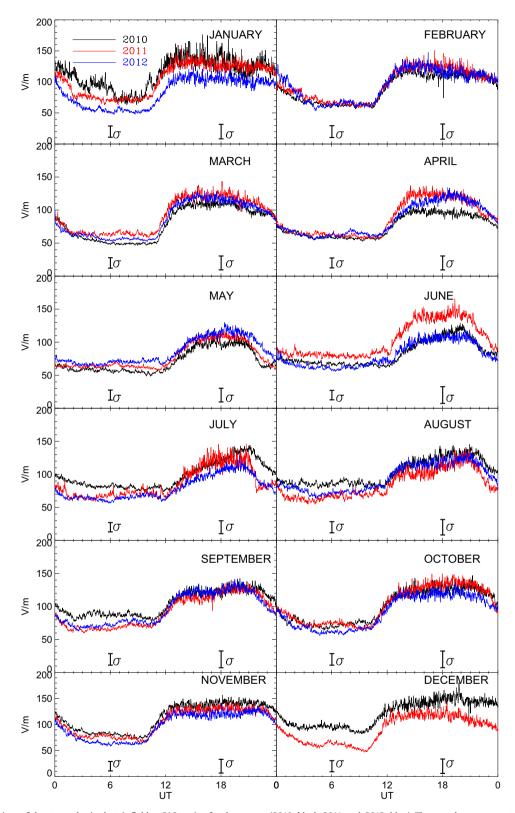


Fig. 7. Monthly variations of the atmospheric electric field at CAS station for three years (2010: black, 2011: red, 2012: blue). The error bars represent 1σ mean values for the three years in the night time (left) and day time (right). (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

existing locations CAS (2 sensors named CAS1 and CAS2 at El Leoncito, San Juan, Argentina), ANC (1 sensor Ancon, Lima, Peru), PLO (1 sensor Punta Lobos, Lima, Peru) and ICA (1 sensor Ica, Peru) are shown by triangles in Fig. 1. New locations for sensor operation

in 2014 are indicated in Fig. 1 by star symbols. A brief description of the stations is indicated in Table 1.

Each sensor consists of a commercially manufactured (Boltek Corporation EFM100-1000120-050205) electric field mill (EFM).

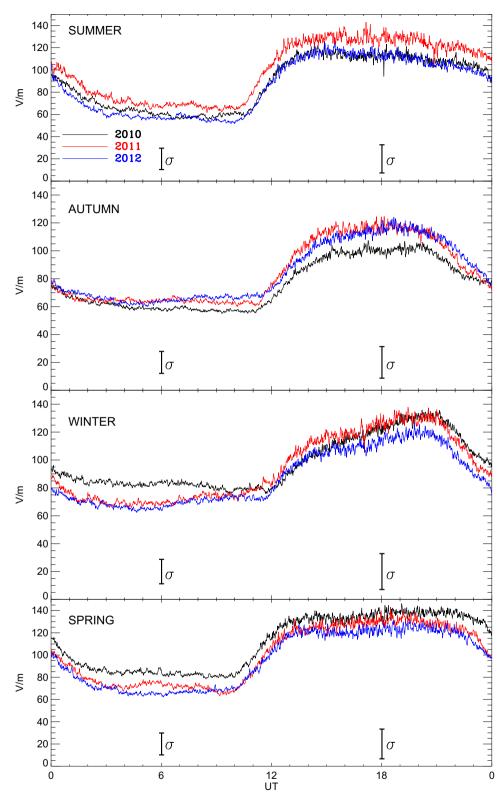


Fig. 8. As in Fig. 7, only that for seasonal variations.

Fig. 2 shows the basic principles of the EFM operation. A current flows from ground through the $100\,k\Omega$ resistor when the sensing electrode is exposed to the atmospheric electric field. When the sensing electrode is shielded from the influence of the electric field, the current flows to ground through the same resistor. The electrical potential drop through the resistor is then proportional to the intensity of the atmospheric electric field. A second resistor

in parallel allows the sensitivity of the measurement to be changed in order to avoid saturation due to high electric field intensities, typically $>\!20~\text{kV/m}$, during stormy or lightning periods.

For actual readings of the atmospheric electric field, the EFM needs to be positioned at ground level. However, this is impractical because the presence of insects, animals, dirt and water can

damage the sensor. For this reason the EFM meter is placed at a certain height above the surface. For example, Fig. 3 illustrates the final position of the EFM sensor located at El Leoncito, at 0.8 m above the ground. However, electric field measurements performed above the ground are abnormally large. This effect is due to an increase of the density of electric field lines and electric charges at a raised conductor. To remove this effect, simultaneous observations have been made with two EFM sensors, named REF and PER. REF and PER are located at ground level and at the final permanent position, respectively. The linear regression between REF and PER recorded data is then calculated, as shown in Fig. 4, and is used to correct daily measurements obtained by the PER sensor. The same treatment is carried out for all the other stations.

In the following, as an example illustrating the network's capability, we show data obtained at CAS2 station. Electric field measurements are collected with a time resolution of 0.5 s, and integrated into 1 min averages for the analysis reported here. The results are summarized in Figs. 5-9. A typical daily curve obtained in fair weather conditions is shown in Fig. 5. Fig. 6 shows the typical diurnal variations during a disturbed day, in this case due to the presence of lightning activity nearby. For comparison, in Figs. 5 and 6, we also show local and Universal Time (LT and UT) in the upper and lower axis, respectively. The methodology adopted in our analysis was to obtain monthly mean curves of the diurnal variations of the atmospheric electric field in fair weather. The criteria for choosing days of fair weather was to select days with potential gradient variations in the range of 40-200 V / m, such as illustrated in Fig. 5. The resulting monthly curves are shown in Fig. 7 for the three years of 2010, 2011 and 2012. The monthly average curves of diurnal variation in fair weather conditions will then be used to compare with daily curves affected by various phenomena mentioned in section 2. Following the same methodology we obtain seasonal and annual curves shown in Fig. 8 and Fig. 9, respectively. The number and proportion of fair weather days selected for the monthly and seasonal curves are summarized in Table 2 and Table 3, respectively. Table 4 shows the mean

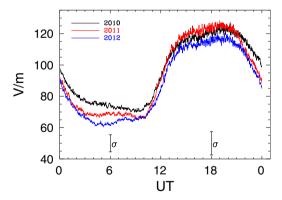


Fig. 9. As in Fig. 7, only that for annual variations.

electrical field value for each hour for the corresponding months in the time period analyzed.

One σ bars are shown in Figs. 7–9, where σ is a mean value of the root mean square deviation obtained from the 3 years of data for 00–12 UT (left) and 12–24 UT (right). It is particularly interesting to note, of these figures, that the shape of the daily variation averaged during a given month, season or year repeats itself almost identically for the different years. Specific time structures of a few hours duration are clearly seen to have similar shapes from year to year.

A gross similarity exists between the annual hourly diurnal variations measured at CAS2 and the Carnegie curve (see, e.g., Fig. 6a from Harrison (2013)). This comparison is shown in Fig. 10. for the three years of 2010, 2011 and 2012. Error bars represent $196\sigma/\sqrt{N}$, where σ is the standard deviation and N is the total number of fair weather days considered. The linear correlation coefficients R are greater than R=0.9 for the three years. In this figure the amplitude values for CAS2 obtained at 2480 masl were multiplied by 1.3 to match the Carnegie curve. This is to account for the exponential decrease of the atmospheric electric field with increasing altitude (Rycroft et al., 2000). However, according to theoretical models, the ratio between Carnegie and CAS2 should be of the order of 2-2.5 rather than 1.3 (Rycroft et al., 2007). One possible explanation for this difference may be because the sensor at CAS2 is located on top of a small hill in a mountainous region, so that the lines of force of the electric field are focused towards the peak, causing an overestimation of the measured electric field values. Indeed, reports of electric field measurements performed at 3400 masl (Cobb, 1968) and 4300 masl (Xu et al., 2013) also show high mean values of 120 V/m and 163 V/m, respectively, higher than what they should at these altitudes.

Another difference when comparing CAS2 measurements and the Carnegie curve concerns the shape of the variation with Universal Time. A great similarity is observed for the time period 14–01 UT with common maxima amplitude around 19–20 UT. However, between 02 and 14 UT we clearly note a large difference. Daily curve measurements performed at land stations are generally different when compared to the Carnegie curve due to local phenomena, such as pollution and aerosols. Similarly, measurement performed at different stations do not show the same minimum and maximum times as in the Carnegie curve

Table 3Total number and proportion of fair weather days selected at CAS station, by season.

| Year/Season | Summer | Autumn | Winter | Spring | Total |
|-------------|--------|--------|--------|--------|-------|
| 2010 | 32 | 51 | 38 | 42 | 163 |
| Proportion | 20% | 31% | 23% | 26% | 100% |
| 2011 | 26 | 48 | 30 | 37 | 141 |
| Proportion | 19% | 34% | 21% | 26% | 100% |
| 2012 | 32 | 53 | 39 | 40 | 164 |
| Proportion | 20% | 32% | 24% | 24% | 100% |

Table 2Total number and proportion of fair weather days selected at CAS station, by month.

| Year/Month | Jan | Feb | Mar | Apr | May | June | July | Aug | Sept | Oct | Nov | Dec | Total |
|------------|-----|-----|-----|-----|-----|------|------|-----|------|-----|-----|-----|-------|
| 2010 | 3 | 17 | 19 | 20 | 13 | 15 | 18 | 10 | 13 | 10 | 18 | 10 | 166 |
| Proportion | 2% | 10% | 11% | 12% | 8% | 9% | 11% | 6% | 8% | 6% | 11% | 6% | 100% |
| 2011 | 9 | 8 | 14 | 16 | 21 | 10 | 4 | 5 | 18 | 12 | 16 | 8 | 141 |
| Proportion | 6% | 6% | 10% | 11% | 15% | 7% | 3% | 4% | 13% | 8% | 11% | 6% | 100% |
| 2012 | 8 | 6 | 21 | 18 | 21 | 8 | 15 | 12 | 20 | 17 | 15 | 0 | 161 |
| Proportion | 5% | 4% | 13% | 11% | 13% | 5% | 9% | 8% | 12% | 11% | 9% | 0% | 100% |

Table 4The mean electrical field value (in V/m) for each hour for the corresponding months for the three years.

| Hour (UT) | January | | | February | | | March | | | April | | | May | | | June | | |
|-----------|------------|------------|----------|------------|------------|----------|-----------|-----------|----------|----------|----------|-----------|----------|----------|----------|----------|-----------|----------|
| | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 |
| 1 | 124 | 109 | 91 | 89 | 92 | 109 | 83 | 83 | 75 | 71 | 76 | 74 | 64 | 64 | 73 | 74 | 86 | 75 |
| 2 | 115 | 94 | 74 | 78 | 86 | 96 | 68 | 70 | 65 | 66 | 68 | 69 | 60 | 64 | 71 | 73 | 82 | 68 |
| 3 | 104 | 77 | 64 | 71 | 83 | 78 | 60 | 65 | 64 | 64 | 64 | 64 | 57 | 63 | 67 | 71 | 83 | 64 |
| 4 | 93 | 77 | 57 | 67 | 74 | 67 | 56 | 64 | 62 | 60 | 63 | 60 | 60 | 64 | 66 | 69 | 82 | 63 |
| 5 | 97 | 74 | 56 | 66 | 72 | 59 | 54 | 62 | 59 | 58 | 65 | 58 | 57 | 63 | 66 | 69 | 81 | 62 |
| 6 | 94 | 70 | 53 | 63 | 66 | 61 | 50 | 64 | 55 | 59 | 63 | 60 | 59 | 64 | 68 | 67 | 81 | 61 |
| 7 | 80 | 70 | 53 | 63 | 66 | 65 | 49 | 65 | 55 | 58 | 62 | 61 | 56 | 68 | 72 | 67 | 79 | 58 |
| 8 | 72 | 72 | 55 | 62 | 65 | 67 | 49 | 66 | 57 | 58 | 61 | 65 | 58 | 64 | 70 | 66 | 79 | 63 |
| 9 | 75 77 | 71 73 | 51 53 | 62 64 | 63 | 64 | 49 49 | 64 | 56 | 57 56 | 61 59 | 67 | 55 54 | 63 | 70 72 | 66 | 78 | 61 65 |
| 10 | | 73 81 | 60 | | 63 | 63 | | 63 | 55 56 | 56 57 | 59 59 | 64 | 54 53 | 62 | 72 73 | 67 | 79 83 | 70 |
| 11 12 | 98 117 | 107 | 77 | 63 83 | 65 88 | 66 89 | 50 64 | 63 80 | 56 72 | 61 | 59 65 | 61 66 | 56 | 61 61 | 73 70 | 66 64 | 83 87 | 70 72 |
| 13 | | | 98 | 83 104 | | 109 | | 80 105 | 97 | 76 | 87 | 79 | 63 | 70 | 76 | 68 | 93 | 72 73 |
| 14 | 133 140 | 123 131 | 106 | 114 | 110 120 | 121 | 88 102 | 119 | 110 | 90 | 107 | 94 | 76 | 70 86 | 87 | 81 | 95 114 | 73 88 |
| 15 | 143 | 133 | 109 | 117 | 125 | 125 | 102 | 122 | 117 | 97 | 122 | 94 106 | 86 | 97 | 101 | 91 | 130 | 97 |
| 16 | 138 | 133 | 109 | 117 | 125 | 123 | 108 | 124 | 121 | 98 | 124 | 109 | 97 | 104 | 101 | 100 | 139 | 103 |
| 17 | 138 | 128 | 107 | 111 | 127 | 124 | 109 | 124 | 118 | 98 97 | 125 | 113 | 97 | 104 | 111 | 104 | 138 | 103 |
| 18 | 137 | 128 | 105 | 111 | 125 | 120 | 108 | 123 | 118 | 98 | 123 | 117 | 101 | 100 | 113 | 107 | 138 | 104 |
| 19 | 137 | 124 | 106 | 112 | 123 | 116 | 110 | 125 | 115 | 96 | 122 | 121 | 96 | 110 | 119 | 114 | 140 | 108 |
| 20 | 125 | 128 | 103 | 113 | 122 | 116 | 108 | 120 | 114 | 94 | 121 | 122 | 99 | 107 | 116 | 117 | 146 | 106 |
| 21 | 117 | 125 | 102 | 111 | 118 | 109 | 104 | 114 | 109 | 95 | 116 | 115 | 98 | 105 | 112 | 116 | 140 | 109 |
| 22 | 123 | 124 | 98 | 113 | 112 | 113 | 97 | 111 | 106 | 91 | 110 | 112 | 85 | 91 | 107 | 101 | 127 | 99 |
| 23 | 131 | 124 | 101 | 109 | 108 | 111 | 96 | 106 | 100 | 88 | 94 | 102 | 65 | 75 | 91 | 86 | 112 | 83 |
| 24 | 121 | 119 | 98 | 101 | 105 | 104 | 93 | 96 | 91 | 79 | 89 | 88 | 66 | 65 | 79 | 83 | 93 | 76 |
| Hour (UT) | July | | | Augus | | | Septen | | | Octobe | | | Novem | | | Decem | | |
| () | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 | 2010 | 2011 | 2012 |
| 1 | 96 | 79 | 71 | 89 | 74 | 84 | 97 | 80 | 81 | 88 | 94 | 88 | 114 | 105 | 98 | 114 | 91 | _ |
| 2 | 90 | 68 | 63 | 86 | 64 | 80 | 88 | 66 | 75 | 78 | 86 | 75 | 99 | 89 | 83 | 100 | 79 | - |
| 3 | 87 | 67 | 66 | 86 | 64 | 78 | 82 | 67 | 67 | 79 | 76 | 68 | 90 | 79 | 73 | 96 | 68 | - |
| 4 | 84 | 64 | 64 | 85 | 60 | 72 | 88 | 65 | 70 | 71 | 72 | 67 | 84 | 76 | 67 | 94 | 62 | - |
| 5 | 81 | 65 | 62 | 84 | 61 | 68 | 91 | 65 | 69 | 70 | 72 | 60 | 83 | 78 | 66 | 97 | 59 | - |
| 6 | 81 | 68 | 59 | 85 | 66 | 70 | 87 | 65 | 72 | 68 | 75 | 61 | 84 | 79 | 62 | 98 | 65 | - |
| 7 | 83 | 68 | 62 | 86 | 67 | 73 | 87 | 67 | 73 | 70 | 75 | 61 | 82 | 74 | 64 | 95 | 63 | - |
| 8 | 80 | 72 | 65 | 90 | 66 | 74 | 88 | 70 | 78 | 73 | 73 | 61 | 79 | 74 | 64 | 95 | 60 | - |
| 9 | 80 | 72 | 66 | 86 | 71 | 73 | 88 | 73 | 79 | 76 | 70 | 64 | 76 | 73 | 64 | 86 | 55 | - |
| 10 | 77 | 75 | 69 | 85 | 68 | 77 | 84 | 71 | 77 | 76 | 68 | 67 | 76 | 69 | 66 | 87 | 50 | - |
| 11 | 80 | 71 | 69 | 84 | 73 | 81 | 81 | 71 | 74 | 79 | 73 | 71 | 92 | 82 | 83 | 100 | 65 | - |
| 12 | 79 | 69 | 63 | 84 | 71 | 83 | 89 | 83 | 81 | 96 | 95 | 86 | 116 | 104 | 101 | 119 | 93 | - |
| 13 | 80 | 72 | 70 | 97 | 88 | 99 | 106 | 97 | 101 | 112 | 111 | 105 | 131 | 121 | 117 | 132 | 110 | - |
| 14 | 90 | 85 | 79 | 109 | 100 | 110 | 114 | 109 | 117 | 120 | 121 | 117 | 138 | 126 | 120 | 137 | 116 | - |
| 15 | 100 | 101 | 88 | 114 | 101 | 116 | 122 | 113 | 121 | 123 | 124 | 117 | 137 | 128 | 121 | 138 | 119 | - |
| 16 | 107 | 108 | 94 | 118 | 101 | 115 | 122 | 111 | 124 | 124 | 128 | 117 | 135 | 128 | 117 | 141 | 120 | - |
| 17 | 115 | 120 | 98 | 120 | 102 | 116 | 121 | 113 | 122 | 124 | 133 | 118 | 138 | 129 | 117 | 142 | 120 | - |
| 18 | 119 | 122 | 102 | 127 | 110 | 117 | 122 | 119 | 121 | 128 | 134 | 120 | 140 | 131 | 120 | 149 | 119 | - |
| 19 | 124 | 119 | 109 | 128 | 113 | 121 | 125 | 126 | 126 | 127 | 135 | 119 | 143 | 132 | 122 | 151 | 120 | - |
| 20 | 133 | 120 | 111 | 131 | 118 | 122 | 129 | 126 | 131 | 127 | 138 | 122 | 142 | 132 | 119 | 154 | 119 | - |
| 21 | 137 | 123 | 111 | 131 | 125 | 125 | 129 | 126 | 128 | 130 | 135 | 119 | 141 | 129 | 124 | 145 | 110 | - |
| 22 | 133 | 101 | 100 | 133 | 115 | 121 | 128 | 124 | 121 | 131 | 132 | 115 | 142 | 129 | 126 | 143 | 106 | - |
| 23 | 117 | 80 | 85 79 | 119 106 | 93 | 109 | 124 | 114 98 | 107 | 122 | 124 | 109 | 139 | 128 | 124 | 145 | 102 | - |
| 24 | 106 | 84 | 79 | 106 | 78 | 93 | 112 | 98 | 92 | 112 | 107 | 99 | 131 | 116 | 111 | 142 | 96 | - |

(Yeboah-Amankwah, 1989; Kamra et al., 1994; Harrison, 2003; Kumar et al., 2009; Guha et al., 2010; De et al., 2013). Nevertheless, in some cases good correlations have been found (Israelsson and Tammet, 2001; Harrison, 2003; Marcz and Harrison, 2003). A possible explanation for the difference found in our results at CAS2 stations, is that measurements are dominated by the effects of tropical convective clouds (Hendon and Woodberry, 1993) and the large lightning rate in South America (Blakeslee et al., 2014). In this case CAS2 measurements show a regional effect rather than a global one. Another possibility would be that the Asian lightning peak occurrence, generally observed in the Carnegie curve around 6-9 UT and assumed to be part of the generator system of the GAEC, does not contribute much to the CAS2 measurements due to attenuation because of the large distance. Indeed, measurements performed over the Indian Ocean show a tendency for higher effects of storm activity over Asia-Australia and Africa-Europe and lower effects of far distant storms over America (Kamra et al., 1994).

Although we have indicated several reasons why our measurements may differ from the Carnegie curves, as indeed other do, to identify the mechanisms for these discrepancies is beyond the scope of this paper and will be studied elsewhere. More important is that we will be able to identify the signatures of possible drivers (see Section 2) in daily time integrated atmospheric electric field variations, if they are enough intense. Therefore, a simple difference between the records obtained during disturbed conditions and the monthly fair weather average, both measured at the same location, will allow isolating the signature of these possible drivers. This should be the procedure in the future, irrespective of whether the average fair weather curves do resemble the Carnegie curve or not, which is a separate research question.

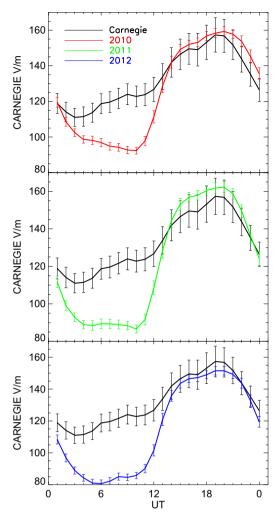


Fig. 10. Annual hourly variations of the atmospheric electric field at CAS station for three years (2010: red, 2011: green, 2012: blue), compared with the Carnegie curve (black). The error bars represent the range of 1.96 standard errors on the mean, corresponding to a 95% confidence interval on the mean. The linear correlation coefficient r between the 24 hourly data values of electric field in CAS and Carnegie are r=0.91 (2010), 0.95 (2011) and 0.94 (2012). (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

4. Conclusions

In this paper we have discussed a new instrumental facility in South America consisting of a network of electric field mills to monitor the vertical atmospheric electric field close to the Earth's surface. Composed of five stations at the present time, we plan to have a few more instruments installed in 2014 which is one the main advantages of this new undertaking in the field of atmospheric electricity. The preliminary results shown here indicate the possibility of providing reliable diurnal variation curves, in all seasons, of the fair weather atmospheric electric field. These curves closely resemble the classical daily variation apparent in the Carnegie curve. Such results will be important to assess how the properties of the GAEC vary as a function of time on different scales. The departures from fair weather diurnal variations will then be used to study the behavior of the GAEC in relation to solar and space weather phenomena and to variations of the galactic cosmic ray flux. We shall also look for short-term atmospheric electricity precursors of enhanced seismic activity in the region, and analyze the observations with respect to local meteorological parameters.

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