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Post-collisional lamprophyric event in Sierra Norte, Córdoba, Argentina: Mineralogical, geochemical and isotopic characteristics

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ABSTRACT

Over 20 lamprophyre dykes, varying in width between a few centimeters and several meters, have been identified in central Sierra Norte – Eastern Pampean Ranges, Córdoba, Argentina. Their mineralogy and chemistry indicate that they are part of the calc-alkaline lamprophyres clan (CAL). They contain phenocrysts of magnesiohornblende ± augite set in a groundmass of magnesiohornblende, calcic-plagioclase, alkali feldspar, and opaque minerals, which designate them as spessartite-type lamprophyres. Alteration products include chlorite, calcite and iron oxides after mafic phenocrysts, though some are partially replaced by actinolite. Feldspars are replaced by carbonate and clay minerals.

The dykes are relatively primitive, and show restricted major element variation (SiO₂ 51.1–55.3 wt.%, Al₂O₃ 12–16.6 wt.%, total alkalies 1.5–4.7 wt.%), high Mg# (55–77), high Cr contents (27–988 ppm) and moderate to high Ni contents (60–190 ppm). Lamprophyre LILE (e.g. Rb averages 110 ppm, Sr 211–387 ppm, Ba 203–452 ppm) are high relative to HFSE (e.g., Ta 0.2–1.6 ppm, Nb 4–11 ppm, Y 17–21 ppm), and are enriched in LREE (30–70 times chondrite). They are characterized by relatively high 208 Pb/ 204 Pb (38.8–39.9), 207 Pb/ 204 Pb (~15.7), and 206 Pb/ 204 Pb (18.7–20.1), combined with low (epsilon) ε_{Nd} (–4.69 to –1.52) and a relative moderately high (87 Sr/ 86 Sr)_i of 0.7055–0.7074. The Rb–Sr whole rock isochron indicates an Early Ordovician age of 485 ± 25 Ma. The calculated T_{DM} (1.7 Ga) suggests that these rocks appear to have originated from a reservoir that was created during a mantle metasomatism event related to the Pampean orogeny.

The Sierra Norte lamprophyres show affinities with a subduction-related magma in an active continental margin. Their geochemical and isotopic features suggest a multicomponent source, composed of enriched mantle material variably contaminated by crustal components. The lamprophyric suite emplacement occurred at the dawning stage of the Pampean orogeny, in a regional post-collisional extensional setting developed in the Sierra Norte-Ambargasta batholith (SNAB) in Early Ordovician times.

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1. Introduction

Lamprophyres are part of an heterogeneous group of porphyritic, mafic-ultramafic, volatile-rich, alkaline rocks, usually forming swarms of dykes, sills or pipe-clusters. Lamprophyres were recently considered as primary melts showing a broad spectrum of compositions in volcanic and plutonic rocks from a variety of geodynamic settings. Calc-alkaline lamprophyre clan constitutes a group of rocks that includes spessartites (hornblende–plagioclase), minettes (phlogopite–K-feldspar), vogesites (amphibole–K-feldspar), and kersantites (phlogopite–plagioclase), and are characteristically associated with plutonic and subvolcanic granitoid rocks

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(Rock, 1987). Previous investigation aimed toward exploring the genetic hypothesis that calc-alkaline lamprophyres originate by contamination of basaltic or lamproitic magmas by continental crust or crustal melts (Rock, 1991).

Lamprophyre dykes are widespread in the Sierras Pampeanas of central Argentina (e.g., Daziano, 1986) but have been the subject of only a few detailed studies, like the Late Famatinian calc-alkaline lamprophyres of the Sierra de San Luis (López de Luchi, 1998; Sparisci and Brodtkorb, 1998; Orozco and Ortiz Suárez, 2005). In the Sierra Norte-Ambargasta Ranges, previous works consist of brief petrographical descriptions of some lamprophyre dykes (Gordillo, 1955; Lucero, 1948, 1969; Sánchez Granel et al., 1967; Minera TEA, 1968; Miró, 2001). Kersantitic lamprophyres have been mentioned in the region between Chuña Huasi and La Puerta (Lucero, 1969) though in this area, only spessartite lamprophyres have been described petrographically.

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In this paper, we present petrographic, mineralogical, geochemical and isotopic information that suggests a complex source for these lamprophyres that, like other calc-alkaline lamprophyres, are associated with a post-collisional extensional event in a subduction-related setting during the final stages of the Pampean orogeny.

2. Geologic setting

The studied lamprophyre outcrops are located north of Córdoba province, between $29^{\circ}28' - 29^{\circ}58'$ S and $64^{\circ}20' - 64^{\circ}00'$ W, in the Sierra Norte-Ambargasta batholith (SNAB), the easternmost block of the Sierras Pampeanas system of Argentina.

In this area the SNAB consists of calc-alkaline, I-type granitoids representative of a Precambrian-Cambrian arc magmatism emplaced in a low grade metasedimentary sequence of Upper Precambrian age. The most primitive rocks of the magmatic sequence are diorite to tonalite microgranular enclaves that are included in the regionally dominant basement formed by granodiorite and monzogranite (GM, Lira et al., 1997) emplaced during Upper Precambrian to Cambrian times (Rapela et al., 1998; Sims et al., 1998; Koukharsky et al., 1999; Millone et al., 2003; Gromet et al., 2005). A differentiated dacitic to rhyolitic subvolcanic body (DR, Lira et al., 1997) called Cerro de los Burros (CDLB) by Lucero Michaut et al. (1995) intruded earlier granitoids; a precise age of this body is in debate, some authors indicate a Late Proterozoic age (Söllner et al., 2000; Millone et al., 2003), others a Late Cambrian age (Miró et al., 2005) and others propose an Early Cambrian age (Leal et al., 2003). Younger intrusions are represented by smaller units of more evolved miarolitic monzogranites of Cambrian age (ca. 525 Ma, Millone et al., 2003), thick tabular syenogranitic aplite bodies (ca. 500 Ma, Millone, 2004) and a considerable number of small Ordovician rhyolitic bodies and dykes (Rapela et al., 1991; Correa, 2003; O'Leary et al., 2006). Although most of these youngest intrusives have ages that, with errors, fall in the Ordovician, geologic setting and petrological and geochemical evidence indicate that they represent the closure of the Pampean orogeny. Only one intrusive event of peraluminous geochemistry, located in the southern part of the batholith, has been considered part of the Famatinian orogeny (i.e., the El Cerro granite; Baldo et al., 1998). A summarized time-event chart of major magmatic and mineralizing episodes in the Sierra Norte-Ambargasta ranges has been recently published by Franchini et al. (2005).

The lamprophyres analyzed in this study occur as dykes that cross-cut both the CDLB body and GM suite (Fig. 1), but more commonly intruded the latter. They have steep dips and randomly oriented strikes, thicknesses range from 0.1 to 5 m and variable lengths from tens to few hundred meters (O'Leary, 2004). These bodies lack chilled margins against their host rocks, and in some cases are associated with thin aplitic dykes.

3. Analytical methods

All samples were crushed to millimeter-scale grain size after removal of weathered rims, only fresh rock chips were selected and this chips were milled in a agate mortar; this material was used for major and trace element analyses which were determined by XRF on fused discs and pressed pellets at XRAL Laboratories (SGS-Canada). REE were determined by ICP-MS at Act-Labs (Canada). Mineral analyses were performed using a CAMECA SX50 electron microprobe in the Department of Geology of Brigham Young University, Provo, Utah, USA. All mineral phases were analyzed using an accelerating voltage of 15 kV, with a beam current of 10 nA and a focused beam of 10 mm in diameter. Natural standards are as follows: Na K α Jadeite, Mg K α MgO, Al K α Anorthite, Si and K K α Ortoclase, Ca K α Wollastonite, Ti K α Titanite, Mn K α Spessartine, Fe K α Fayalite. For micro-analytic data, reduction was done based on the correction factors of Pouchou and Pichoir (1985).

The isotopic analyses were carried out at the Geochronological Research Center (CPGeo) of the Institute of Geosciences, University of São Paulo, Brazil. All samples were crushed to millimeter-scale grain size after removal of weathered rims and hand-picked under a stereoscopic magnifier. These chips were washed in purified water in an ultrasonic bath. Powered samples were used for Rb-Sr, Pb-Pb, and Sm-Nd determinations. The Pb-Pb analyses were made by acid digestion and Pb purification by carbonate ionic exchange using the NBS981 standard. The Rb-Sr and Sm-Nd analyses were prepared by standard methods according to the analytical procedures described by Kawashitamodified (1972), and Sato et al. (1995), involving HF-HNO3 dissolution plus HCl cation exchange. Analyzed samples had no visible solid residues after dissolution. The Sr isotopic ratios were normalized to 86 Sr/ 88 Sr = 0.1194: replicate analyses of ⁸⁷Sr/⁸⁶Sr for the NBS987 standard gave a mean value of 0.71028 ± 0.00006 (2σ), the blanks for Sr were 5 ng. Nd ratios were normalized to a 146 Nd/ 144 Nd = 0.72190. The averages of ¹⁴³Nd/¹⁴⁴Nd for La Jolla and BCR-1 standards were 0.511847 ± $0.00005(2\sigma)$ and $0.512662 \pm 0.00005(2\sigma)$, respectively. The blanks were less than 0.03 ng. Pb, Sr, and Nd isotopic analyses were carried out on a multicollector VG 354 Micromass and Finnigan-MAT 262 mass spectrometers. The isotopic data were regressed using the program of Ludwig (1999). The ages were calculated using the decay constants established in Steiger and Jäger (1977); for ⁸⁷Rb was 1.42×10^{-11} yr⁻¹; for ¹⁴⁷Sm = $6.54 \times$ 10^{-12} yr^{-1} .

4. Petrology of the Sierra Norte lamprophyres

4.1. Petrography and mineralogy

The Sierra Norte lamprophyres are mesocratic to melanocratic rocks, hypidiomorphic, fine grained and microporphyritic to porphyritic with an aphanitic groundmass. In some cases banded structures indicating flow differentiation are present. Euhedral to subhedral phenocrysts include calcic amphibole, in some cases clinopyroxene and, more rarely, plagioclase; phenocrysts occur as individual grains and as radiating clusters forming a glomeroporphyritic texture. The groundmass consists of calcic amphibole, in addition to microlites of plagioclase and fine-grained potassium feldspar; both quartz and plagioclase occur as intergranular phases. Accessory minerals set in the groundmass include needles of apatite and opaque minerals (magnetite–hematite, bornite + chalcopyrite + chalcocite–covellite, and pyrite).

Many of the dykes are extensively altered (deuterically) and the primary minerals are largely replaced. Alteration products include chlorite, carbonate, epidote, and iron oxides after calcic amphiboles; clinopyroxenes and calcic amphiboles are occasionally transformed to a secondary amphibole, and feldspars are altered to clay minerals and sericite, epidote and carbonate. Locally, some dykes exhibit ocellar texture; ocelli are spheroidal to ovoidal aggregates, consisting of quartz and calcite, and are interpreted as globular structures mineralogically analogous to basaltic amygdales (Rock, 1984). Xenocrystic quartz and feldspars occur in some samples, suggesting partial crustal contamination by assimilation of country rock.

Petrographic characteristics of the dykes are generally consistent with their inclusion in the lamprophyre family of rocks. In terms of the modal mineralogy, the dykes are spessartites (plagioclase, hornblende, augite) according to the IUGS classification (Streckeisen, 1979; Wooley et al., 1996).

Clinopyroxenes were identified in only one dyke (1391) and are classified as augite (Morimoto et al., 1988) with high TiO_2 and low



Fig. 1. Geological map of central Sierra Norte-Ambargasta in the eastern block of the Pampean Ranges of Argentina, showing the main rock types. A and B display the local geology of the sampled lamprophyres (they were enlarged for an easier identification).

NaO₂. Representative analyses of mafic silicates are included in Table 1. Augites are in the compositional range of pyroxenes of calcalkaline lamprophyres (CAL; Rock, 1987), and show an inverse relation between Mg# and the TiO₂ and Al₂O₃ contents. The relation between Al^{IV} and Al^{VI} contents in pyroxenes is used as an indicator of pressure conditions; in this case, the lamprophyric augites have a composition of high pressure pyroxenes, with Al^{IV} and Al^{VI} contents that range from 0.07 to 0.08 and 0.04 to 0.15, respectively (Bédard et al., 1988; Dorais, 1990).

The composition of phenocrysts of amphibole have little variation from one sample to another (structural formulae were calculated following the I.M.A. indications, Leake et al., 1997; Table 1); the lamprophyre dykes contain magnesiohornblende with very uniform composition, only one dyke contains edenite (sample 1399). The microprobe analyses indicate that the amphiboles are homogeneous and show no evidence of zoning or reaction with the melt. The cationic diagram Ca + Al^{IV} vs. Si + Na + K shows that the lamprophyric amphiboles present a compositional trend, starting with edenitic cores in sample 1399 and hornblende rims in samples 1391–1399–1413, extending to actinolitic hornblendes in sample 1306 (Fig. 2). This evolutionary trend is distinguished by decreasing Al, Ti, Ca, Na and K, increasing Si, Fe and Mn, and almost constant Mg, Fe and Ti-tschermakitic and edenitic-type of substitutions (O'Leary, 2004), which suggest that the hornblendes, and probably the actinolitic hornblendes as well, are products of magmatic crystallization.

Additional evidence of primary equilibrium crystallization of amphiboles from the melt is from computed Mg# for the amphiboles using Bédard's (1988) equation (Mg# amphibole = 0.253125 ^{*}Mg# whole rock + 0.5425). The calculated values correspond closely to actual Mg#s of the amphiboles in the studied spessartites (Table 1). However, secondary SiO₂-poor actinolitic amphiboles after both primary amphiboles and clinopyroxenes were recognized.

Plagioclase from groundmass is relatively anorthitic (An_{56-76}). Plagioclase in sample 1306 shows normal zoning with less calcic rims (An_{41-49}), whereas sample 1391 plagioclase exhibits an inverse zoning with andesine–oligoclase core and labradorite rims.

Table 1

Representative electron microprobe compositions of mafic silicates of spessartite lamprophyres from central Sierra Norte-Ambargasta

	1306	1306	1306	1306	1306	1306	1399	1399	1399	1399	MUE1	MUE1	MUE1	1413	1413
	1	1	2	2	1	2	1	1	2	2	1	1	3	1	2
	Core	Rim	Core	Rim	Altered	Altered	Core	Rim	Core	Rim	Core	Rim	Intergr. cpx	Core	Core
Amphibo	oles														
SiO ₂	50.06	49.54	51.35	50.12	51.66	52.31	47.04	46.98	43.27	43.73	47.54	49.96	53.16	50.55	46.65
TiO ₂	0.08	0.17	0.29	0.20	0.10	0.09	0.92	0.95	2.33	2.27	1.66	0.94	0.50	0.29	1.37
AI_2O_3	5.49	5.85	4.31	5.54	4.09	3.75	7.23	6.91	9.63	9.78	7.41	5.37	3.25	4.35	6.59
MnO	0.40	0.41	0.46	0.48	039	0.37	0.27	0.31	0.05	0.22	0.29	039	0.34	0.48	0.58
MgO	13.75	14.38	15.28	14.93	15.14	15.94	15.09	14.63	14.30	14.36	15.12	15.26	17.14	13.39	11.19
CaO	12.36	12.14	11.75	11.65	12.56	12.36	10.70	10.89	11.03	11.10	11.12	10.53	11.81	11.77	11.48
Na ₂ O	0.40	0.48	0.38	0.43	0.31	0.29	1.14	1.08	1.70	1.85	1.43	0.98	0.43	0.46	0.77
K ₂ O	0.14	0.20	0.23	0.48	0.12	0.19	0.34	0.34	0.61	0.60	0.45	0.29	0.16	0.28	0.45
r CI	0.26	0.20	0.17	0.20	0.19	0.21	0.09	0.14	0.50	0.27	0.21	0.19	0.12	0.30	0.28
Total	97.28	96.79	97.50	97.03	97.05	97.07	93.92	93.62	94.07	95.15	98.16	97.02	97.04	97.39	97.99
F + CI	0.11	0.09	0.09	0.09	0.08	0.09	0.06	0.08	0.14	0.13	0.11	0.10	0.06	0.34	0.35
Total	97.17	96.70	97.41	96.94	96.97	96.98	93.87	93.55	93.93	95.03	98.06	96.92	96.98	97.05	97.64
Structur	al formulae	based on () = 23												
^{IV} Si	7.448	7.369	7.430	7.288	7.599	7.552	7.008	7.054	6.534	6.536	6.891	7.262	7.614	7.411	6.932
"Al Tsite	0.552	0.631 8.000	0.570	0.712	0.401	0.448	0.992	0.946	1.466	1.464	1.109	0.738	0.386	0.589	1.068
^{VI} AI	0.120	0.103	0.166	0.237	0.094	0.164	0.278	0.276	0.247	0.258	0.157	0.182	0.163	0.163	0.089
^{vI} Tl	0.008	0.019	0.031	0.022	0.011	0.010	0.103	0.108	0.264	0.255	0.181	0.103	0.054	0.032	0.154
Fe ³⁺	0.306	0.383	0.299	0.320	0.195	0.241	0.417	0.332	0.306	0.285	0.396	0.295	0.085	0.326	0.496
Mg	3.050	3.189	3.295	3.236	3.321	3.430	3.350	3.274	3.220	3.199	3.196	3.306	3.661	2.926	2.478
re" Mn	1.480	1.285	1.209	1.183	1.341	1.146	0.852	1.010	0.962	1.003	1.069	1.114	1.037	1.553	1.782
C site	5 000	5 000	5.000	5 000	4 998	5 000	5.000	5 000	5.000	5 000	5.000	5.000	5 000	5 000	5 000
Fe ²⁺	0.000	0.000	0.091	0.073	0.000	0.005	0.106	0.080	0.074	0.076	0.094	0.175	0.087	0.019	0.032
Mn	0.014	0.031	0.056	0.058	0.012	0.037	0.033	0.039	0.027	0.028	0.035	0.048	0.042	0.059	0.072
Ca	1.970	1.934	1.799	1.815	1.974	1.912	1.709	1.753	1.784	1.778	1.726	1.640	1.812	1.849	1.829
Na Risto	0.016	0.035	0.054	0.054	0.014	0.042	0.152	0.128	0.115	0.118	0.145	0.137	0.059	0.073	0.067
Ca	0.000	2.000	2.000	0.000	2.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
Na	0.099	0.103	0.054	0.075	0.075	0.039	0.178	0.186	0.382	0.416	0.256	0.140	0.060	0.057	0.156
K	0.026	0.038	0.042	0.090	0.022	0.034	0.064	0.065	0.117	0.114	0.083	0.053	0.029	0.052	0.085
A site	0.125	0.141	0.096	0.165	0.103	0.074	0.243	0.250	0.499	0.530	0.339	0.193	0.089	0.110	0.241
Mg#" Mg#b	63.1 70.7	65.7	68.6	68.2	68.4	71.2	72.5	70.9	71.7	71.3	68.6 73.0	70.1	76.5	61 68	51.7
Pyroxen	es				MU	E1	73.7			MU	E1			00	MUE1
-					1					2					2
										(a)					(h)
6:0					52.0	25				(a)	1				(0)
SIO_2 TiO ₂					52.0	J5 7				52.0)]				52.17
Al ₂ O ₂					2.6	, 7				3.29	r)				3.40
FeO					5.82	2				5.80)				6.47
MnO					0.10	5				0.28	3				0.17
MgO					16.5	50				16.2	26				16.41
Na ₂ O					0.19	24 J				21.4 0.16	10 5				017
K ₂ O					0.01	1				0.02					0.01
Total					99.0	02				99.8	35				100.56
Structur	al formulae	based on () = 4												
Si					1.92	22				1.92	21				1.926
Al					0.07	78				0.07	'9				0.074
Fe T site					2.00	00				2.00	0				2 000
Al					2.00	38				2.00	54				2.000
Fe					0.10	08				0.00	79				0.199
Ti					0.0	1				0.01	5				0.016
Cr															
Mg					0.84	44				0.74	12				0.711
M1 site					1.00	00				1.00	00				1 000
Mg					0.00	65				0.15	54				0.193
Mn					0.00	05				0.00	8				0.005
C-					0.84	40				0.84	19				0.838
Cd										0.01					0012
Ca Na K					0.0	14				0.01	1				0.012
Ca Na K M2 site					0.0	14 24				0.00	2)1 24				1.048
Ca Na K M2 site Wo					0.0 0.92	14 24 72				0.00 0.00 1.02 0.45	2 01 24 52				1.048 0.441
Ca Na K M2 site Wo En					0.92 0.92 0.22	14 24 72 53				0.01 0.02 1.02 0.45 0.48	2 01 24 52 89				1.048 0.441 0.466
Ca Na K M2 site Wo En Fs					0.0 0.9 0.2 0.5 0.1	14 24 72 53 75				0.01 0.02 0.45 0.48 0.05	2 1 2 4 5 2 3 9 5 8				1.048 0.441 0.466 0.093

^a Mg/(Mg + FE² + Fe³ ^b Bédard (1988).

Table 2

Si0₂

Ti0₂

 Al_2O_3

Sample wt.%



Fig. 2. Cationic diagram Al^{IV} + Ca vs. Si + Na + K showing a compositional trend of amphiboles from Sierra Norte spessartite lamprophyres. The black stars indicate the end-members compositions.

Potassium feldspars were identified in all samples but only are analyzed in 1391 groundmass and have a constant composition (Or_{96}). Feldspar compositions are shown in Fig. 3.

4.2. Whole-rock geochemistry

Results of whole-rock major and minor element analyses, REE analyses and normative compositions are presented in Table 2. The lamprophyres are chemically equivalent to subalkaline basalts in the TAS diagram (Fig. 4a). Analytical data were recalculated on a volatile-free basis and show low SiO₂ contents (51.1–55.3 wt.%), Al₂O₃ (12–16.6 wt.%), CaO (4.2–8.7 wt.%), and moderate total alkali contents (1.5–4.7 wt.%) are somewhat lower than those of andesites and basalts from orogenic regions (e.g., Suzuki and Shiraki, 1980; Wilson, 1989). They are calc-alkaline lamprophyres according to the classification scheme of Rock (1987) in a K₂O vs. SiO₂ diagram (Fig. 4b). The dykes are over-saturated with respect to silica, containing both normative hypersthene and quartz. In terms of fer-



Fig. 3. Feldspars analyses of Sierra Norte spessartite lamprophyres plotted in the Or-Ab-An triangle of Deer et al. (1963).

e ₂ O ₃	10.00	8.54	7.87	8.19	8.17
ligO InO	7.95	0.15	0.13	0.15	4.58
	8.94	7.00	4.51	7.42	7.89
a0 la-0	1 32	2.19	4.51	1.48	2.96
-0	1.52	1 39	0.92	0.91	1 29
205	0.13	0.13	0.16	0.13	0.20
r ₂ O ₃	0.06	0.17	0.03	0.19	n.d.
Ō	2.20	3.30	6.60	3.15	1.81
otal	100.10	100.13	100.25	100.01	99.99
1g# ^a	65	74	63	77	55
pm					
b	123	53	34	23	105
r	213	239	250	211	387
a	262	369	452	267	203
	21	19	18	17	21
Г Ib	110	109	126	113 6	134
10	4	0 <2	o s d	s d	<2
σ	<0.5	<0.5	s.d.	s.d.	<0.5
-5 1	<0.5	<0.2	s.d.	s.d.	<0.5
n	2	1	s.d.	s.d.	2
b	1.4	<0.5	s.d.	s.d.	2.3
S	11.3	6	s.d.	s.d.	14.9
f	3	2.9	s.d.	s.d.	4.1
a	0.2	0.4	s.d.	s.d.	0.3
V	3	20	s.d.	s.d.	2
1	1	0.5	s.d.	s.d.	0.8
b	<5	6	s.d.	s.d.	<5
i	<0.4	<0.4	s.d.	s.d.	<0.4
h	1.8	4	s.d.	s.d.	4.1
	0.4	0.9	s.d.	s.d.	1.4
	213	166	s.d.	S.G.	179
r o	335	988	139	814 ad	27
0	40	40	s.u.	s.a.	<22
11 11	134	49	s.u. s.d	s.u. s.d	<10
n	98	45 87	s.u. s.d	s.u.	83
a	17	15	s.d.	s.d.	19
e	2	2	s.d.	s.d.	3
s	<5	<5	s.d.	s.d.	6
с	s.d.	33	s.d.	s.d.	25
		1206	1201		1412
ample ppm		1300	1591		1415
a		9.80	13.70)	20.50
e		20.40	27.90)	42.20
r		2.65	3.34	_	4.92
ld		11.90	13.80)	19.90
m		3.00	3.00		4.10
u 'd		1.10	0.88		1.40
h		0.60	0.50		0.60
UV IV		3 90	3.00		3.80
0		0.80	0.60		0.70
r		2.40	1.80		2.20
m		0.34	0.26	0.32	
b		2.20	1.70		2.20
u		0.36	0.27		0.34
EE total		62.73	73.75	107.04	
REE		47.75	61.74	91.62	
IREE		13.80	11.13	3	13.96
a/Lu		27.22	50.74	1	60.29
IPW normativ					
ample	e compositions	1001	1207	1000	
ampie	e compositions 1306	1391	1395	1399a	1413
l	e compositions 1306 9.06	1391 7.26	1395 9.60	1399a 10.96	1413
r	e compositions 1306 9.06 10.28	1391 7.26 8.51	1395 9.60 5.81	1399a 10.96 5.57	1413 12.35 7.77
r b	e compositions 1306 9.06 10.28 11.40	1391 7.26 8.51 19.15	1395 9.60 5.81 37.65	1399a 10.96 5.57 12.94	1413 12.35 7.77 25.48
r b n	e compositions 1306 9.06 10.28 11.40 31.99	1391 7.26 8.51 19.15 22.81	1395 9.60 5.81 37.65 20.50	1399a 10.96 5.57 12.94 25.59	1413 12.35 7.77 25.48 29.70
r b n li (wo)	e compositions 1306 9.06 10.28 11.40 31.99 5.25	1391 7.26 8.51 19.15 22.81 5.15	1395 9.60 5.81 37.65 20.50 1.01	1399a 10.96 5.57 12.94 25.59 4.89	1413 12.35 7.77 25.48 29.70 3.75
r b n li (wo) li (en)	e compositions 1306 9.06 10.28 11.40 31.99 5.25 4.53	1391 7.26 8.51 19.15 22.81 5.15 4.44	1395 9.60 5.81 37.65 20.50 1.01 0.87 0.87	1399a 10.96 5.57 12.94 25.59 4.89 4.22	1413 12.35 7.77 25.48 29.70 3.75 3.23
e r b n vi (wo) vi (en) vi (fs)	e compositions 1306 9.06 10.28 11.40 31.99 5.25 4.53 0.00	1391 7.26 8.51 19.15 22.81 5.15 4.44 0.00	1395 9.60 5.81 37.65 20.50 1.01 0.87 0.00	1399a 10.96 5.57 12.94 25.59 4.89 4.22 0.00 26.60	1413 12.35 7.77 25.48 29.70 3.75 3.23 0.00
r b n ii (wo) ii (en) ii (fs) ly (en)	e compositions 1306 9.06 10.28 11.40 31.99 5.25 4.53 0.00 15.73 0.00	1391 7.26 8.51 19.15 22.81 5.15 4.44 0.00 22.71 2.20	1395 9.60 5.81 37.65 20.50 1.01 0.87 0.00 14.81 0.00	1399a 10.96 5.57 12.94 25.59 4.89 4.22 0.00 26.30 2.00	1413 12.35 7.77 25.48 29.70 3.75 3.23 0.00 7.92 6 22
r b n vi (wo) vi (en) vi (fs) ly (en) ly (fs)	e compositions 1306 9.06 10.28 11.40 31.99 5.25 4.53 0.00 15.73 0.00 2.67	1391 7.26 8.51 19.15 22.81 5.15 4.44 0.00 22.71 0.00 2.51	1395 9.60 5.81 37.65 20.50 1.01 0.87 0.00 14.81 0.00 0.45	1399a 10.96 5.57 12.94 25.59 4.89 4.22 0.00 26.30 0.00 0.5	1413 12.35 7.77 25.48 29.70 3.75 3.23 0.00 7.92 0.00 0.00
r b n i (wo) i (en) i (fs) ly (en) ly (fs) nt	e compositions 1306 9.06 10.28 11.40 31.99 5.25 4.53 0.00 15.73 0.00 0.67 9.76	1391 7.26 8.51 19.15 22.81 5.15 4.44 0.00 22.71 0.00 0.51 8.40	1395 9.60 5.81 37.65 20.50 1.01 0.87 0.00 14.81 0.00 0.45 8.00	1399a 10.96 5.57 12.94 25.59 4.89 4.22 0.00 26.30 0.00 0.51 8.12	1413 12.35 7.77 25.48 29.70 3.75 3.23 0.00 7.92 0.00 0.62 7.90
r b n i (wo) i (en) i (fs) iy (en) iy (fs) nt e	e compositions 1306 9.06 10.28 11.40 31.99 5.25 4.53 0.00 15.73 0.00 0.67 9.76 0.00	1391 7.26 8.51 19.15 22.81 5.15 4.44 0.00 22.71 0.00 0.51 8.49 0.00	1395 9.60 5.81 37.65 20.50 1.01 0.87 0.00 14.81 0.00 0.45 8.09 0.00	1399a 10.96 5.57 12.94 25.59 4.89 4.22 0.00 26.30 0.00 0.51 8.12 0.00	1413 12.35 7.77 25.48 29.70 3.75 3.23 0.00 7.92 0.00 0.62 7.89 0.00
r r i (wo) i (fs) y (en) y (fs) tt e	e compositions 1306 9.06 10.28 11.40 31.99 5.25 4.53 0.00 15.73 0.00 0.67 9.76 0.00 0.29	1391 7.26 8.51 19.15 22.81 5.15 4.44 0.00 22.71 0.00 0.51 8.49 0.00 0.29	1395 9.60 5.81 37.65 20.50 1.01 0.87 0.00 14.81 0.00 0.45 8.09 0.00 0.37	1399a 10.96 5.57 12.94 25.59 4.89 4.22 0.00 26.30 0.00 0.51 8.12 0.00 0.29	1413 12.35 7.77 25.48 29.70 3.75 3.23 0.00 7.92 0.00 0.62 7.89 0.00 0.62 7.89 0.00 0.62

Major, minor, and trace elements concentrations and normative compositions of

1395

54.30

0.79

14.90

1399

53.50

12.50

0.59

some spessartite lamprophyres from central Sierra Norte-Ambargasta

1306

51.10

1.02

15.50

1391

52.90

0.66

13.20

*0.899 = FeO.

^a (Mg/Mg + Fe)100-Fe₂O₃.

1413

55.32

0.81

16.97



Fig. 4. Chemical classification of the lamprophyres. (a) TAS diagram (dates are re-calculated to a volatile-free basis). The line separates alkaline from subalkaline rocks (Irvine and Baragar, 1971). (b) Classificatory scheme of lamprophyres (Rock, 1987), samples as in Fig. 4a. The studied lamprophyres are typical medium-K calc-alkaline type.

romagnesian element abundances, the high Mg# (55–77), Cr (27– 988 ppm) and Ni contents (60–190 ppm) indicate the relatively primitive nature of the dykes (Rock et al., 1988).

The lamprophyres show an enrichment of Rb (23–147 ppm), Ba (203–452 ppm), Sr (211–387 ppm) and other large ion lithophile elements (LILE) relative to HFSE (high field strength elements), especially Ta (0.2–1.6 ppm), Nb (4–11 ppm) and Y (17–21 ppm). MORB-normalized spidergrams (Fig. 5) show an irregular pattern, characterized by relative enrichments (15–120 times) of the

incompatible elements (Rb, K and Ba), with a strong positive spike in Rb and negative troughs in Nb, Ta, P and Ti, whereas Sr shows a rather uniform normalized concentration clustered around $Sr_N = 3$. The Cs_N is 850 to 2200 times higher than MORB. This pattern is similar to those of the average spessartite (Rock, 1991) and other calc-alkaline lamprophyres (e.g., Tate and Clarke, 1993; Hoch et al., 2001; Guo et al., 2004). In comparison, the studied spessartites patterns are slightly depleted in Sr, Th, Ta, Ce and P (HFSE and LREE) and present a positive spike in Rb. The Sierra Norte lam-



Fig. 5. MORB normalized multi-element diagram comparing the geochemical compositions of studied lamprophyres with average spessartite (Rock, 1991) and Phanerozoic continental margin basalts (Pearce, 1983). Normalization values after Sun and McDonough (1989). The inset shows the chondrite normalized REE diagram (normalization factors after Boynton, 1989). See text for further details.

Table 3	
Sr. Nd. and Pb isotopic compositions of spessartite lamprophyres from central Sierra Norte-Ambargasta	

Sample	Rb (ppm)) Sr (ppr	n) ⁸⁷ Rb	^{/86} Sr	Error	⁸⁷ Sr/ ⁸⁶ Sr	Error	(⁸⁷ Sr/ ⁸⁶ Sr) ₀	$\varepsilon_{Sr}(T)$
1391	57.5	244.5	0.68	1	0.006	0.712047	0.000035	0.70738	49.18
1399	31.6	220.2	0.41	5	0.014	0.708352	0.000081	0.70550	22.49
1413	110.6 395.0		0.811		0.006	0.712947	0.000052	0.70738	49.26
1306	136.5	224.8	1.76)	0.026	0.719504	0.000061	0.70743	49.94
	Sm (ppm)	Nd (ppm)	¹⁴⁷ Sm/ ¹⁴⁴ Nd	Error	143Nd/144Nd	Error	$(^{143}Nd/^{144}Nd)_0$	$\varepsilon_{\rm Nd}$ (T)	$T_{\rm DM}~({\rm Ga})$
1391	3.183	13.964	0.1378	0.0005	0.512211	0.000009	0.511773	-4.31	1.7
1399	3.114	13.766	0.1368	0.0005	0.512241	0.000008	0.511806	-3.66	1.6
1413	4.275	19.821	0.1304	0.0004	0.512196	0.000008	0.511782	-4.14	1.5
1306	3.123	12.002	0.1573	0.0005	0.512435	0.000012	0.511935	-1.14	1.6
	²⁰⁶ Pb/ ²⁰⁴ Pb		Error (2σ)	207	Pb/ ²⁰⁴ Pb	Error (2σ) ²⁰⁸		²⁰⁴ Pb	Error (2σ)
1391	19.014	1	0.009	15	.671	0.007	39.173		0.013
1399	19.207		0.005	15.701		0.005	39.477		0.005
1413	20.137		0.015	15.759		0.015	39.908		0.015
1306	6 18.749		0.012	.012 15.692		0.014	38.860		0.013

prophyres seem to be broadly similar in chemical composition to modern and Phanerozoic continental margin basalts (Pearce, 1983); Fig. 5 shows a similar pattern with an average composition of basalts except for a larger enrichment of K and Rb.

The transition elements (V, Cr, Co and Ni) show different behaviors according to the degree of differentiation; the most primitive samples have compositions close to those of MORB, particularly the sample 1391 which has high Cr abundances (988 ppm), whereas the most differentiated sample (1413) is depleted in the same elements respect to MORB.

Chondrite normalized data (Fig. 5) exhibit fractionated REE distribution patterns (Boynton, 1989) with enrichment for LREE (30– 70 times), and a flat pattern of HREE (<10 times chondrite). Samples 1306 and 1413 show a slightly positive Eu anomaly and sample 1391 shows a smooth negative Eu anomaly (Eu/Eu^{*} = 0.9). The more primitive Sierra Norte lamprophyres are depleted in REE respect to other lamprophyric suites (e.g., Tate and Clarke, 1993; Hoch et al., 2001; Guo et al., 2004), except for sample 1413, which represents the most geochemically evolved lamprophyre of the studied samples.

4.3. Sr, Nd, and Pb isotopes

Sr, Nd, and Pb isotope compositions were studied to provide further constrains on the source and crustal interaction during magma evolution. The analytical results are listed in Table 3.

The present-day ratios of ⁸⁷Sr/⁸⁶Sr from Sierra Norte lamprophyres range from 0.70835 to 0.71950, and ¹⁴³Nd/¹⁴⁴Nd range from 0.51219 to 0.51243. Within the Rb-Sr isochronic diagram three analytical points (whole-rock samples 1391, 1306 and 1413) defined a line with a slope that corresponds to an age of 485 ± 25 Ma, with an initial strontium ratio of 0.706 ± 0.016 (Fig. 6). The (epsilon) ε_{Sr} at 485 Ma ranges from 49.9 to 22.5, and individual (⁸⁷Sr/⁸⁶Sr)_i values are between 0.70743 and 0.7055. All lamprophyres have negative (epsilon) ϵ_{Nd} values calculated for 485 Ma, ranging from -4.31 to -1.14, with a relatively constant 143 Nd/ 144 Nd, between 0.51193 and 0.51177. In the (87 Sr/ 86 Sr)_i vs. (epsilon) $\epsilon_{Nd(485 Ma)}$ diagram (Fig. 7) the samples plot between EM I and EM II (enriched mantle type I and II). In comparison, these spessartites are less radiogenic than calc-alkaline lamprophyres of other localities (e.g., minettes, Hoch et al., 2001) but Late Carboniferous lamprophyres from Europe show similar isotope compositions. The European lamprophyres (minettes and kersantites) are interpreted as melts of a mantle source that was enriched due to the recycling of continental crust at an active margin (Hegner et al., 1998; Turpin et al., 1988). The Pb isotope compositions are ²⁰⁸Pb/²⁰⁴Pb (38.8–39.9), ²⁰⁷Pb/²⁰⁴Pb (~15.7), and ²⁰⁶Pb/²⁰⁴Pb



(18.7–20.1). In the ²⁰⁸Pb/²⁰⁴Pb and ²⁰⁷Pb/²⁰⁴Pb vs. ²⁰⁶Pb/²⁰⁴Pb dia-

grams the samples plot close the upper crust curve of the plumbo-

Fig. 6. Whole rock Rb–Sr isochron of Sierra Norte-Ambargasta spessartite lamprophyres. (Data-point error ellipses are 68.3% conf.). Samples plotted are 1391, 1413 and 1306.



Fig. 7. $(^{87}Sr)^{86}Sr)_i$ vs. ϵ_{Nd} diagram. The contoured fields contain published data of calc-alkaline lamprophyres from other localities. BE: bulk earth. EM I and II: enriched mantle type I and II (Zindler and Hart, 1986). See text for further details.

tectonic model of Zartman and Doenot shown (1981), and fall near the mantle Northern Hemisphere Reference Line of Hart (1984), close to lamprophyric rocks of other localities (Fig. 8).

The Sm–Nd depleted-mantle model ages (DePaolo, 1981) are around 1.7 Ga, what suggest the age of the mantle metasomatism event associated with magmatic arc evolution during the Pampean orogeny.

In summary, the available isotopic information of studied lamprophyres indicate enriched mantle-derived magmas as the main magma's source, although some crustal contamination during the magma ascent to the upper crustal levels can not be rejected.

5. Discussion

5.1. Petrogenesis

The Sierra Norte lamprophyres can be classified as spessartites; considering the phenocrysts mineralogy, they were recognized as augite spessartites and hornblende spessartites. Petrographical, mineralogical, geochemical and isotopic features of dykes indicate a common genetic relationship and allow us to propose a similar origin and a common parental magma.

The presence of minerals with elongate habit and porphyritic textures suggests rapid crystallization and cooling during the emplacement. Xenocrystic quartz and feldspars occur in some samples, suggesting partial crustal contamination by assimilation of country rock. Clinopyroxenes composition suggests a high crystallization pressure and is compatible with the presence of associated hornblende.

Suzuki and Shiraki (1980) discussed the stability range of the hornblende–clinopyroxene assemblage in experimental studies of water-saturated systems ($SiO_2 = 50-60$ wt.%), concluding, with reference to crystallization pressure, that: (a) the phenocrystic hornblendes in hornblende–spessartites might have crystallized on the liquidus from a magma of andesitic composition at pressures greater than 8 Kbar under water-saturated conditions, and (b) the groundmass hornblendes and the fine-grained spessartites might have crystallized under pressures between 4 and 1 kbar.

The similarities of the whole-rock SiO₂ content and the composition of the phenocrysts of the early mafic assemblage of the studied Sierra Norte lamprophyres, with the experimental data of Suzuki and Shiraki (1980), might suggest that mafic phenocrystic phases crystallized under initial pressures that were much higher than final emplacement pressures. This fact would also imply fast ascent and cooling to prevent mafic phenocrysts from reequilibration.

The emplacement pressure for the lamprophyres is estimated to be less than 3 kbar; this estimation is based on the fact that lamprophyric dikes crosscut brittle dacite–rhyolite rocks of the subvolcanic CDLB body and of the granodiorite–monzogranite country rocks of the GM suite, which, according to Poklepovic et al. (2005) were emplaced at maximum pressures of 2.5 ± 0.5 kbar and 3 ± 0.6 kbar, respectively.

The slight compositional variation of hornblende in all lamprophyres, the absence of hornblende zonation, the Mg# in agreement with Bédard's Mg# value, and its chemical composition indicate a primary origin for these amphiboles (Pe-Piper, 1988). The principal substitutions registered in lamprophyric hornblendes suggest crystallization from a mafic to intermediate silica-saturated to oversaturated alkaline magma.

Experimental evidence in basalts indicate that the stability field of plagioclase is reduced by increasing water pressure (Bédard, 1988); some previous contributions (e.g. Lange et al., 1993; Carmichael et al., 1996) consider that in magmas with SiO₂ concentrations close to 60 wt.%, the abundance of primary amphibole and the nearly absent phenocrystic plagioclase could be evidence of melts with high water contents. In the same direction, experimental work that aimed to reproduce hornblende phenocrysts without coexisting plagioclase or pyroxene phenocrysts in spessartites, needed water saturation at \sim 3 kbar, corresponding to >5.5% water in the melt (Moore and Carmichael, 1998). At lower pressures and concentrations of water, hornblende was replaced by orthopyroxene, augite and plagioclase (Carmichael et al., 1996).

5.2. Magma source

Most of the dykes are representative of primitive compositions, with high Mg# and Ni and Cr contents close to MORB, and show high concentrations of transition elements. In some cases (samples 1306 and 1413), the slightly positive Eu anomaly restricts the lamprophyric magma formation to a free-plagioclase source, (e.g. peridotite or pyroxenite), and in others (sample 1391), the smooth negative Eu anomaly (Eu/Eu^{*} = 0.9) can be explained by minor plagioclase fractionation. The HREE contents, <10 times chondrite, are compatible with retention of these elements in the residual magma as a result of fractionation of garnet in the source, and the HFSE content range is similar to that of a depleted mantle. The LILE abundance (15–220 times MORB) and the LREE (30–70 times chon-



Fig. 8. Pb isotope diagrams of Sierra Norte spessartite lamprophyres compared with MORB and lamprophyres from other localities. See text for further details.

drite) suggest an incompatible elements enriched source. The Ba/ La (9–27), Pb/La (0.4–1.9) and Cs/Rb (0.09–0.15) ratios are very high compared to MORB values (3–10, 0.09–0.12 and 0.013, respectively), which suggest an important contribution from sediments to magma, also consistent with the Ce positive anomaly (Ce/ Ce^{*} = 2.15–2.63). The (⁸⁷Sr/⁸⁶Sr)_i, (epsilon) ε_{Nd} , and high ²⁰⁷Pb/²⁰⁴Pb and high ²⁰⁶Pb/²⁰⁴Pb values indicate a mantle-derived magma with an important crustal contamination. The origin of a crustal component in lamprophyric magmas has been studied by many authors (e.g., Turpin et al., 1988; Carmichael et al., 1996; Hoch et al., 2001), from whom two classical interpretations arise: (a) contamination of mantle-derived magma during emplacement, and (b) formation of an enriched lithospheric mantle reservoir during a subduction process. Our data is insufficient to be able to favor one interpretation over the other.

Several authors (e.g. MacDonald et al., 1985; Barnes et al., 1986) propose a multiple origin source hypotheses for some calc-alkaline lamprophyres, resulting from a contribution of sub-continental lithospheric mantle (LILE high contents and HFSE variably high) and astenospheric mantle (HFSE, Cr and Ni high contents) with a metasomatic component derived from subducted pelagic sediments (low Nb contents and high volatiles) and variably contaminated by a crustal component. The Sierra Norte lamprophyres show a geochemical and isotopic signature compatible with those characteristics, indicating similar sources.

5.3. Tectonic setting and emplacement age

The calc-alkaline lamprophyres are, in essence, products of subduction-related magmatism (Rock, 1991). They are usually emplaced at convergent margins as dykes that intruded calc-alkaline granitic rocks. Several authors related lamprophyres to continental collision events, whereas others indicated its relation with arc setting; in both, the cortical material, crust or sediments, are recycled into the mantle (Hoch et al., 2001). Using some immobile elements (Ta/Yb vs. Th/Yb diagram, Pearce, 1983, not shown) the studied rocks of Sierra Norte suggest an active continental margin setting, and derivation from an enriched mantle metasomatized during the subduction of oceanic lithosphere. Other evidence for this source is the Zr and Y contents (109–120 ppm and 17–25 ppm, respectively) which relate the Sierra Norte lamprophyres with active continental margin magmas (Pearce, 1983). Likewise, the high Ba/Ti values are typical of subduction zone magmas. The ages of lamprophyres (see below) and their geologic setting within the Sierra Norte magmatic arc relate their intrusion to the terminal stage of subduction.

The isotopic features indicate a possible derivation from an enriched subcontinental mantle source, modified by a subduction event, and petrography suggests that this magma was contaminated with upper continental crust during ascent and/or emplacement. This geological setting was previously proposed as the origin of the Sierra Norte-Ambargasta batholith (Ramos, 1988; Kraemer et al., 1995; Lira et al., 1997; Rapela et al., 1998), indicating the emplacement of a magmatic arc in the protomargin of Gondwana as a consequence of the approach of the Pampia terrane to Córdoba basement, with the main magmatic activity taking place during the Late Proterozoic to the Late Cambrian (Leal et al., 2003; Millone et al., 2003). Recently, another scenario was proposed (Gromet et al., 2005, and references therein) who suggest a ridge subduction process without the existence of a continental collision as origin for the Sierra Norte-Ambargasta batholith.

The isotopic signature of the lamprophyres is compatible with a post-collisional emplacement during the Ordovician times, and the T_{DM} (1.7 Ga) suggests the involvement of a crust with affinities with a basement located to the west of a remarkable tectonic boundary existent between the Rio de la Plata craton and the Pampean Ranges in agreement with the most recent publications (Leal

et al., 2003; Rapela et al., 2005; Gromet et al., 2005). The age of the lamprophyres (485 ± 25 Ma) is virtually identical to the emplacement ages of other small bodies and dykes from the Oncán rhyolite unity (Rapela et al., 1991; Correa, 2003; Millone, 2004; O'Leary et al., 2006) and is comparable with the age of a hydrothermal event registered in the Cerro de los Burros dacite–rhyolite subvolcanic body (Leal et al., 2003). All these ages support the existence of a regional post-collisional event during the Early Ordovician, ca. 490 Ma.

6. Conclusions

Calc-alkaline lamprophyre dykes emplaced in the Sierra Norte-Ambargasta batholith during Early Ordovician times are spessartites; their main mineralogy is represented by magnesiohornblende \pm augite and calcic plagioclase, which is product of primary magmatic crystallization. The Sierra Norte lamprophyric suite geochemically resembles basalts and low silica andesites, with SiO₂ contents ranging from 51.1 to 55.3 wt.%, moderate Na₂O + K₂O contents from 1.5 to 4.7 wt.%, and high volatile contents (LOI values ranging from 1.8 to 6.6). The lamprophyres have Mg# between 55 and 77, and high Cr contents (27–988 ppm).

The calc-alkaline nature of Sierra Norte lamprophyres is depicted in the MORB-normalized patterns, which show an enrichment in LILE relative to HFSE, and a negative Nb–Ta anomaly that is typical of calc-alkaline lamprophyres characteristic of destructive plate-margin magmatism; this has been interpreted as indicative of assimilation of subducted sediment. Chondrite-normalized REE patterns are fractionated with Ce_N = 25–52 and Yb_N = 8.1–10.5. The lamprophyres show a limited range in initial ⁸⁷Sr/⁸⁶Sr (0.705–0.707), (epsilon) $\varepsilon_{Nd(T)}$ (–1.45 to –4.57), and have high ²⁰⁷Pb/²⁰⁴Pb and high ²⁰⁶Pb/²⁰⁴Pb values. The isotopic features indicate a possible derivation from an enriched subcontinental mantle source, modified by a subduction event, with subordinated contributions of upper crust material to lamprophyric magma. The T_{DM} (1.7 Ga) suggests a crustal component with affinities to the basement of the Pampean Ranges.

The mineralogical, geochemical and isotopic information indicates a source related to a subduction setting and suggest the emplacement of the Sierra Norte lamprophyric suite at the final stage of the Pampean orogeny, in a regional extensional setting developed in the Sierra Norte-Ambargasta batholith during a post-collisional event in the Early Ordovician.

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